

The study of the $N = 28$ shell closure : a way to probe nuclear forces.

O. Sorlin¹

¹*Grand Accélérateur National d'Ions Lourds (GANIL),
CEA/DSM - CNRS/IN2P3, B.P. 55027, F-14076 Caen Cedex 5, France*

The study of the evolution of the $N = 28$ magic shell has been started about 15 years ago. A lot of experimental studies aimed at determining whether the $N = 28$ shell closure is eroded in very neutron-rich nuclei through complementary methods. The very first hints of the vanishing of the $N = 28$ gap were obtained by β -decay [1], mass measurements [2] and Coulomb-excitation [3,4] experiments. Theoretical studies came progressively to the conclusion that the erosion of the $N = 28$ shell gap should lead to shape coexistence in ^{44}S and deformation in ^{42}Si (e.g.). For this latter, it was however not clear whether spherical shape would be preferred to deformation. Recent experimental campaigns were carried out worldwide to study atomic masses [5], β -decay [6], nuclear spectroscopy [7,8,9,10,11,12], neutron single-particle energies [13] to probe the onset of collectivity along the $N=28$ isotones. With these pieces of information in hand and new unpublished results, the progressive collapse of the $N = 28$ shell closure is now established. Interestingly this study is also ideal for probing the nuclear forces such as the spin-orbit and tensor forces in nuclei. Experimental highlights will be shown at the $N = 28$ shell closure, in connection with the state of the art theoretical descriptions. More generally, the shell-breaking mechanism discovered at $N=28$ should apply to other shell closures, such as the $N = 14$, $N = 50$ and $N = 82$ ones, in which the same spin orbit forces are at play. However they seem to have various consequences with respect to shell erosion...

- [1] O. Sorlin et al., Phys. Rev. C. 47 (1993) 2941
- [2] F. Sarazin et al., Phys. Rev. Lett. 84 (2000) 5062
- [3] H. Scheit et al, Phys. Rev. Lett. 77 (1996) 3967
- [4] T. Glasmacher et al., Phys. Lett. B 395 (1997) 163
- [5] B. Jurado et al., Phys. Lett. B. 649 (2007) 43
- [6] S. Grévy et al., Phys. Lett. B 594 252 (2004)
- [7] D. Sohler et al., Phys. Rev. C 66 (2002)054302
- [8] P.M. Campbell et al., Phys. Rev. Lett. 97 (2006) 112501
- [9] S. Grévy et al, Eur. Phys. J. A 25 (2005) s01-111
- [10] B. Bastin et al., Phys. Rev. Lett. 99 (2007) 022503
- [11] S. Bhattacharyya et al., Phys. Rev. Lett. 101 (2008) 032501
- [12] L. Gaudefroy et al., Phys. Rev. Lett. 102 (2009) 092501
- [13] L. Gaudefroy et al., Phys. Rev. Lett. 97 (2006) 092501