



Opinion: Beyond global means – novel space-based approaches to indirectly constrain the concentrations of and trends and variations in the tropospheric hydroxyl radical (OH)

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Abstract. The hydroxyl radical (OH) plays a central role in tropospheric chemistry, as well as influencing the lifetimes of some greenhouse gases. Because of limitations in our ability to observe OH, we have historically relied on indirect methods to constrain its concentrations, trends, and variations but only as annual global or annual semi-hemispheric averages. Recent methods demonstrated the feasibility of indirectly constraining tropospheric OH on finer spatio-temporal scales using satellite observations as proxies for the photochemical drivers of OH (e.g., nitrogen dioxide, formaldehyde, isoprene, water vapor, ozone). We found that there are currently reasonable satellite proxies to constrain up to about 75 % of the global sources of tropospheric OH and up to about 50 % of the global sinks. With additional research and investment in observing various volatile organic compounds, there is potential to constrain an additional 10 % of the global sources and 30 % of the global sinks. We propose steps forward for the development of a comprehensive space-based observing strategy, which will improve our ability to indirectly constrain OH on much finer spatio-temporal scales than previously achieved. We discuss the strengths and limitations of such an observing strategy and potential improvements to current satellite instrument observing capabilities that would enable better constraint of OH. Suborbital observations (i.e., data collected from non-satellite platforms such as aircraft, balloons, and buildings) are required to collect information difficult to obtain from space and for validation of satellite-based OH estimates; therefore, they should be an integral part of a comprehensive observing strategy.

1 Introduction

The hydroxyl radical (OH) has a pivotal role in atmospheric chemistry and controls the lifetimes of methane (CH₄), hydrochlorofluorocarbons and hydrofluorocarbons (HCFCs and HFCs), carbon monoxide (CO), volatile organic compounds (VOCs), and numerous other gases relevant to climate and air quality. The relationship of OH to other atmospheric constituents is often complex and non-linear. For instance, the concentrations and lifetimes of ozone (O₃) and OH are convolved, as the primary source of OH involves the destruction of O₃, OH is integral in many tropospheric chemical reactions (e.g., initiating the oxidation of VOCs) that lead to O₃ production, and the reaction of HO₂ with NO is a secondary source of OH that can also lead to O₃ production. The reader is referred to, for instance, Lelieveld et al. (2016) and Fiore et al. (2024) for comprehensive overviews of the importance of OH in tropospheric chemistry, as well as discussions of key uncertainties in its global chemical budgets and estimates of past global trends and variations in its atmospheric abundance.

In situ measurements provide information about tropospheric OH at a particular point in time and space. However, these observations are sparse, and OH is highly variable in space and time as its lifetime is < 1 s. Therefore, in situ observational networks are not a viable way to map the variability and trends in OH across the globe. In addition, there are few in situ instruments that can quantify OH at typical ambient concentrations of 10⁵–10⁷ molec. cm⁻³ (e.g., Stone et al., 2012). Generally, in situ measurement uncertainties are ~ 30 % for integration times of 30–60 s (Brune et al., 2018). Comparing OH observations from multiple airborne missions to steady-state box model simulations, Miller and Brune (2022) concluded that “oxidation chemistry in most of the free troposphere is understood to as well as current measurements can determine.” This implies that for some parts of the atmosphere, further refinements in our understanding of tropospheric oxidation are limited by inherent measurement uncertainties for both OH and related species. Nevertheless, the suite of other trace gases and aerosols observed during campaigns, both with and without the sparse in situ OH data, have provided a detailed, albeit geospatially limited, view of tropospheric composition (e.g., Nicely et al., 2016).

A tropospheric OH data product derived from a spaceborne instrument is not currently feasible, although it may become feasible with technological development. The total column of OH has been measured for decades during day-time with passive instruments at a few ground-based facilities (e.g., Cageao et al., 2001; Minschwaner et al., 2003), and new technology development could lead to an OH lidar (Pan et al., 2022). Studies would need to be performed to assess the suitability of these passive and active instruments for deployment on satellites. If it is feasible to deploy such instruments in space, a tropospheric column of OH could be inferred by subtracting the stratospheric column observations

from the total column observations; space-based instruments that measure stratospheric-mesospheric OH, albeit not down to the tropopause, have already been demonstrated (e.g., the Microwave Limb Sounder, MLS; Wang et al., 2008).

There are a number of methods to indirectly constrain tropospheric OH on global scales, such as with suborbital observations (e.g., Lelieveld et al., 2006). Methods that use in situ data of methylchloroform or a combination of HCFC observations constrain only global or, possibly, hemispheric OH (e.g., Liang et al., 2017; Thompson et al., 2024). For methylchloroform, recent declines in tropospheric abundance limit its utility (Lelieveld et al., 2004; Liang et al., 2017). Nevertheless, multi-decadal methylchloroform observations have provided an important constraint on the long-term trends and variations in tropospheric OH (e.g., Patra et al., 2021, and references therein). Shortwave infrared (SWIR) observations of CH₄ from the Japan Aerospace Exploration Agency (JAXA) Greenhouse gases Observing SATellite (GOSAT) have been proposed as a way to indirectly constrain global OH (Zhang et al., 2018), although additional surface and thermal infrared (TIR) observations may be required (Zhang et al., 2021).

In addition to the lack of data for spatio-temporal constraint, uncertainties in the atmospheric processes (e.g., trace gas and heterogeneous chemistry, emissions) that influence tropospheric OH (e.g., as summarized in Table 2 of Fiore et al., 2024; Prather and Zhu, 2024) have resulted in a large range in simulated OH among chemical transport models (CTMs; e.g., Nicely et al., 2017, 2020; Zhao et al., 2019; Murray et al., 2021). Consequently, there is considerable uncertainty in the sink of CH₄, which contributes significantly to the overall uncertainty in the budget of CH₄ (Saunio et al., 2020).

Recent efforts have demonstrated satellite-based methods to indirectly constrain the spatio-temporal concentrations, trends, and variations in OH on local and regional scales (e.g., Valin et al., 2013, 2016; Wolfe et al., 2019; Wells et al., 2020; Pimlott et al., 2022; Zhu et al., 2022b; Anderson et al., 2023, 2024; Zhao et al., 2023; Souri et al., 2024; Shutter et al., 2024; Zhu et al., 2024). These methods use various satellite datasets as proxies for OH drivers (e.g., CH₄, water vapor (H₂O_(v)), nitrogen dioxide (NO₂), CO, formaldehyde (HCHO), isoprene, tropospheric O₃). Thanks to contemporary multi-decadal data records, these approaches open new avenues for characterizing OH over long periods of time, which is important for understanding the impacts of both anthropogenic activities and natural phenomena on tropospheric OH.

The purpose of this paper is to make recommendations for developing a strategic and comprehensive observing strategy, which will improve our ability to indirectly constrain OH on much finer spatio-temporal scales than has previously been achieved. Our recommendations are intended to inform international efforts to prioritize observational needs. In Sect. 2, we describe the current space-based approaches to constrain

tropospheric OH. In Sect. 3, we summarize suborbital observations (i.e., data collected from non-satellite platforms, such as aircraft, balloons, and buildings) that are required for information currently unobtainable from space (e.g., fine vertical resolution of satellite proxies) and for validation of satellite-based OH estimates. In Sect. 4, we discuss potential refinements to retrieval algorithms as well as to instrument capabilities that would improve satellite data products. Finally, in Sect. 5, we discuss potential improvements to current satellite observing capabilities that would better enable constraint of OH. These improvements include ones that are obtainable with current technologies (e.g., more observations, co-located observations) as well as ones requiring additional technology development (e.g., to obtain finer vertically resolved observations).

2 Current approaches to indirectly constrain local and regional concentrations of and trends and variations in tropospheric OH with satellite observations

Here we provide a brief overview of the various space-based approaches that have been used to constrain tropospheric OH as partial and/or total vertical column densities (VCDs) or surface concentrations on monthly to seasonal timescales. Each of these approaches has strengths and limitations that need to be considered when being applied to constrain tropospheric OH.

2.1 Process-based approaches

These approaches exploit photochemical relationships between satellite-observable species and tropospheric OH, which aid understanding of the causes of trends and variations in inferred OH (e.g., Valin et al., 2016; Nicely et al., 2018; Baublitz et al., 2023).

- Wolfe et al. (2019) found a linear relationship between OH and HCHO using in situ observations from NASA's Atmospheric Tomography (ATom) field campaign, which is consistent with the modeling work of Valin et al. (2016), who found that HCHO VCDs primarily depend on OH production rates at low OH concentrations. This indicates that HCHO VCDs, such as from the Ozone Monitoring Instrument (OMI) on NASA's Aura satellite, could be used to infer tropospheric OH VCDs on local scales. This approach is valid over remote oceanic environments, where CH₄ oxidation is the primary HCHO source; there are broad regions of the troposphere that are characterized by this photochemical environment.
- Pimlott et al. (2022) developed an approach that employs a simplified steady-state approximation with multiple trace gas observations (i.e., O₃, CO, CH₄, and

H₂O_(v)) from the Infrared Atmospheric Sounding Interferometer (IASI) on the European Space Agency (ESA) MetOp-A satellite. This approach is most representative of OH from 600 to 700 hPa, although the authors argue that their approach could be reasonably applied from 400 to 800 hPa.

- Shutter et al. (2024) developed an approach to constrain tropospheric OH trends and variability that uses the ratio of VCDs of isoprene from the NOAA Suomi NPP Cross-track Infrared Sounder (CrIS) and HCHO from the Ozone Mapping Profiler Suite (OMPS) over isoprene-emitting forest ecosystems. They take advantage of the fact that the ratio of the VCDs of isoprene and HCHO scales with 1/OH in such environments (Wells et al., 2020).

2.2 Machine learning (ML) approaches

These approaches use a combination of an ML model, output from a CTM, and satellite observations. The utility of these approaches for constraining tropospheric OH depends on the accuracy of the training dataset. For example, there are known knowledge gaps in our understanding of the processes that influence OH (Fiore et al., 2024). In addition, these approaches reproduce OH only for photochemical environments that are represented in the training dataset. For example, the following two approaches were designed for polluted urban environments and cleaner tropical and subtropical ones, respectively.

- Zhu et al. (2022b) developed an approach that takes advantage of the known response of surface OH to NO_x (= NO₂ + NO) concentrations to infer trends and variations in summertime surface OH over selected cities. Their six input parameters, on which their ML model was trained, explain 76 % of the variance in simulated surface OH. Two of the input parameters are satellite data (i.e., OMI NO₂ and HCHO), although the other parameters are simulated by a CTM.
- Anderson et al. (2023) developed a multi-satellite approach, testing the utility of datasets of numerous trace gases (i.e., Table 1 of Anderson et al., 2023) that influence OH through chemical, radiative and meteorological effects. They demonstrated the potential of their approach for inferring trends of tropospheric OH VCDs over tropical and subtropical oceans (Fig. 1), and their approach is being expanded over land and the extratropics. Preliminary results indicate that this approach, which uses tropospheric and total VCDs, may be used to infer OH for various layers of the troposphere; information on the vertical distribution of OH is important for estimating, for instance, the lifetime of CH₄, given that CH₄ loss is weighted toward the lower troposphere.

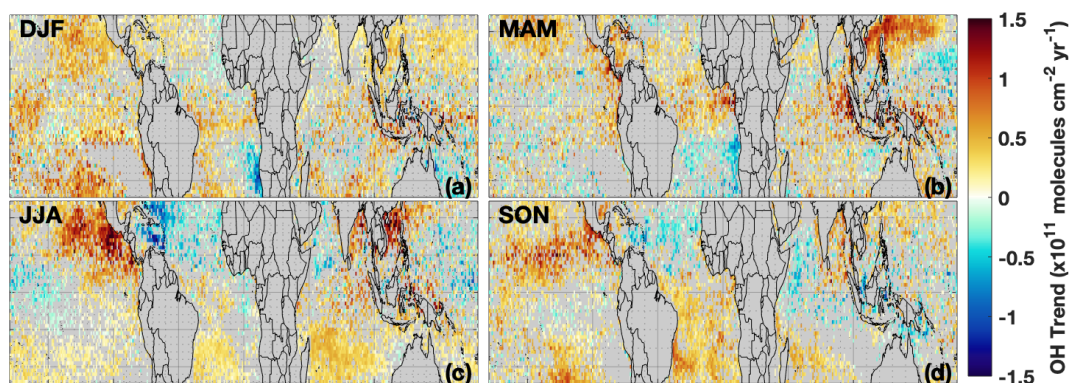


Figure 1. There is significant seasonal and spatial variability in the trends in the tropospheric VCD of OH over tropical oceans from 2005 to 2019. We calculated the trends using the OH product described in Anderson et al. (2023) and a multiple linear regression described in Anderson et al. (2024). Colored areas indicate those grid boxes where the trend is statistically significant and with at least 10 years of data. This figure is an adaptation of Fig. 2 from Anderson et al. (2024).

2.3 Chemical data assimilation approaches

Historically, these approaches incorporate satellite observations of OH drivers into CTMs to constrain tropospheric OH. They use inverse modeling and/or chemical data assimilation methods (Bocquet et al., 2015) to reconcile modeled estimates with observations while relying on the underlying chemical mechanisms of the model to infer OH concentrations consistent with the assimilated observations. A notable benefit of data assimilation is that it provides an improved estimate of concentrations, and the resulting increments indicate where model parameterizations are deficient and/or require improvement. Overall, the benefit of data assimilation depends on the magnitude of the difference between the observed and simulated variable.

- Miyazaki et al. (2012, 2020) and several other papers from the same group developed multi-species (notably NO_2 , O_3 , and CO) assimilation systems for estimating tropospheric composition and emissions. In particular, assimilation reduced the multi-model spread and annual biases of key OH drivers in four CTM frameworks. The multi-model spread for tropospheric mean OH was reduced by 24 %–58 % over polluted areas.
- Gaubert et al. (2016, 2017) assimilated Measurements of Pollution in the Troposphere (MOPITT) CO observations into the Community Atmosphere Model with Chemistry (CAM-Chem) of the Community Earth System Model (CESM). They showed, for example, that increases in CO because of assimilation can lead to decreases in OH and finally to longer lifetimes of CH_4 . In contrast, decreases in CO can lead to increased OH and shorter CH_4 lifetimes. Highlighting nonlinear effects, the study also demonstrated that assimilation led to changes in O_3 and photochemical activity due to HO_x recycling.

2.4 Simplified hybrid approaches

Several recent simplified approaches allow for the quantification of the impacts of constraining various OH drivers with satellite observations, which could inform decision-making (e.g., prioritization of observables) in inverse model and data assimilation efforts (Sect. 2.3).

- Zhao et al. (2023) developed a method to post-process a CTM simulation to improve simulated tropospheric OH. Their method adjusts the simulated monthly averaged concentrations of OH drivers (CO , CH_4 , O_3 , HCHO , NO_2 , total column O_3 , and $\text{H}_2\text{O}_{(\text{v})}$) with satellite observations for each model column and then uses a photochemical box model to recalculate the corresponding concentration of tropospheric OH.
- Zhu et al. (2024) used an ML technique to develop an emulator of simulated tropospheric OH VCDs in the CESM Whole Atmosphere Community Climate Model (WACCM), which facilitated their assessment of the roles of satellite VCDs of NO_2 , HCHO , and CO on tropospheric OH concentrations, trends, and variations.
- Souri et al. (2024) developed a Bayesian data fusion method that adjusts monthly averaged distributions of OH drivers from a CTM simulation using satellite observations. This method considers the quality of the datasets and the a priori used in the retrieval. The adjusted model fields are then used in a parameterization of OH (Anderson et al., 2022) in a CH_4 – CO –OH cycle model (ECCOH; Elshorbany et al., 2016), allowing for easy evaluation of the impact of adjustments on OH, CH_4 , and CO .

The robustness of constraining OH with these satellite-based approaches is not limited by, for instance, the ML or data assimilation methods at the present time but instead by the quality of the data used as input to these methods

and limitations in our current understanding of the chemistry and emissions that influence tropospheric OH (e.g., Table 2 of Fiore et al., 2024). Therefore, the recommendations in the sections that follow focus on further development of the satellite-based approaches as well as improvement to the quality of satellite data and to independent validation products.

3 Recommendations for suborbital needs

The utility of any space-based OH estimate (Sect. 2) hinges on an adequate understanding of how OH responds to its drivers. Therefore, long-term measurements of key atmospheric properties, supplemented with occasional air- and ground-based intensive campaigns, would build confidence far beyond what has been achieved with previous intermittent efforts, such as the NASA Atmospheric Tomography Mission (AToM) that provided a comprehensive suite of observations, primarily in remote oceanic environments. In addition, a comprehensive suite of co-located observations, both in situ and remotely sensed, is desired for validating CTMs, developing process-based diagnostics (Sect. 5.3), interpreting trends and variations in inferred OH, and improving satellite retrievals (Sect. 4).

We recommend the development of a long-term, robust suborbital observing network composed of instruments at stationary sites as well as ones deployed in field campaigns. The establishment of OH “supersite” observatories would allow for the collection of a comprehensive suite of co-located measurements of key atmospheric constituents (discussed below) to constrain tropospheric OH production and loss (Sect. 5.2.1). Focusing first on augmenting existing efforts (e.g., NOAA; In-service Aircraft for a Global Observing System, IAGOS; Network for the Detection of Atmospheric Composition Change, NDACC; and Advanced Global Atmospheric Gases Experiment, AGAGE) would keep early-development efforts feasible by leveraging existing facility infrastructure. In addition, some existing efforts (e.g., four NOAA baseline observatories, the IAGOS program) already observe many of the variables that reasonably constrain OH, as discussed below. Periodic field campaigns would supplement the data collected at OH supersites, especially vertical composition profiles and in photochemical environments not represented by the ground and routine aircraft networks. Continuous and co-located surface and VCD observations would be ideal, especially over a range of photochemical environments (e.g., remote to urban) and seasons. Planetary boundary layer height data would be useful for interpreting these observations. The placement and density of OH supersites and locations/timing of field campaigns would benefit from observing system simulation experiments (OSSEs) in order to maximize the value of the suborbital data for constraining OH. For instance, a simpler strategy would likely be satisfactory over the relatively homogeneous remote oceans

(e.g., fewer stations and observations of individual VOCs) as compared to more complex land environments. Therefore, the OSSE studies would also need to consider the required suite of observations at individual sites.

We recommend a focus on the collection of suborbital observations of OH drivers, given the lower uncertainties typically associated with these measurements as compared to those for OH (Sect. 1). Such observations would allow for the constraint of OH within photochemical box models, although OH budget closure remains an active area of research (Stone et al., 2012), and it is unclear whether existing in situ observations can adequately explain OH in all cases (e.g., Fuchs et al., 2017; Lew et al., 2020; Hansen et al., 2021; Yang et al., 2022; Bottorff et al., 2023; Yang et al., 2024).

The following variables reasonably constrain tropospheric OH in photochemical box models and should be priority measurements: $\text{H}_2\text{O}_{(\text{v})}$, NO, NO_2 , CO, O_3 , actinic flux (or filter radiometers for wavelength ranges relevant to NO_2 and $\text{O}(^1\text{D})$), HCHO, and temperature. In regions closer to strong emissions sources (over and near land), chemical controls on OH are more complex and variable and a more comprehensive measurement suite (e.g., total OH reactivity, VOC composition, nitrous acid (HONO), and peroxy radicals (HO_x , RO_x)) (e.g., Lelieveld et al., 2016; Yang et al., 2016; Murray et al., 2021) may be needed to constrain all drivers. In remote marine environments, information on VOC speciation and reactivity is less critical but not unimportant (Travis et al., 2020; Baublitz et al., 2023), as the background atmosphere is (relatively) uniform, and large-scale OH variability primarily reflects variability in OH production from photolysis and HO_x cycling (Wolfe et al., 2019; Brune et al., 2020).

Occasional deployment of in situ instruments that observe tropospheric OH at OH supersite observatories and in field campaigns would allow for an assessment of the consistency between observed OH and OH inferred from the satellite-based approaches (Sect. 2). As mentioned above, there are few in situ instruments, and technology development is necessary before such observations could become routine.

4 Recommendations for retrieval algorithm refinement

Here, we discuss potential refinements to retrieval algorithms, as well as to instrument capabilities, that would improve data products for the satellite-based approaches (Sect. 2).

4.1 Ultraviolet/visible (UV/Vis) data

The quality of UV/Vis satellite-based data products is determined by a number of issues, including signal-to-noise ratio (SNR), spectral fitting uncertainties, systematic and random measurement errors, and uncertainties associated with creating VCDs (e.g., Duncan et al., 2014; Lorente et al., 2017). For example, to extract the spectral signature of

HCHO used in retrievals, several interfering geophysical processes need to be accounted for, including absorption by O₃, NO₂, bromine monoxide (BrO), and oxygen dimers (O₂–O₂), as well as rotational Raman scattering (Ring effect; e.g., González Abad et al., 2016). Coupled with its small atmospheric loading, especially over remote oceanic areas, the HCHO VCDs are “noisy”, requiring temporal and/or spatial averaging to reduce random errors so that the geophysical signal becomes clear (e.g., Liao et al., 2024). As another example, NO₂ VCDs are dominated by the stratospheric component over remote areas, leading to relatively large uncertainties in the corresponding tropospheric portion of VCDs (e.g., Lamsal et al., 2021). Improvements in data quality, whether through data product refinement (below), oversampling techniques, or enhancements in instrument capabilities (Sect. 5), will be required to use the satellite-based approaches (Sect. 2) at finer spatial and temporal resolutions than are achievable with current instruments.

Work is ongoing to improve the accuracy and precision of satellite data products through retrieval algorithm refinement and also using, for example, data-driven ML data analysis methods that show promise for reducing noise in data products when the SNR is low and/or the trace gas is a weak absorber (e.g., Li et al., 2022; Joiner et al., 2023). For instance, Joiner et al. (2024) have tested their technique on NASA Tropospheric Emissions: Monitoring of Pollution (TEMPO) HCHO retrievals. Preliminary results (Fig. 2) show visibly less noise while keeping overall consistency with the original noisy data, and there is a reduction in the number of pixels with negative slant column density (SCD) values in the noise-reduced data.

Additionally, different retrieval methods applied to the same instrument often produce large differences in data products, which impact inferred OH (e.g., Anderson et al., 2023). Discrepancies created by retrieval algorithm assumptions could be resolved with more robust suborbital data than is currently available for validation (Sect. 3), especially for a priori profiles used in retrievals.

Additional potential areas of retrieval algorithm refinement include aerosol correction and stratosphere–troposphere separation that will benefit from improvements in radiative transfer forward model simulations as implemented in level 2 (L2) retrieval algorithms. For example, to reduce computation time, many retrieval algorithms use pre-computed lookup tables and do not explicitly represent important geophysical processes such as aerosol scattering/absorption. With sufficient acceleration, radiative transfer models or ML proxies can be used directly in L2 algorithms, reducing interpolation errors and providing greater flexibility to use input data that are more representative of the atmospheric state (e.g., the aerosol profiles from MERRA-2 reanalysis used in OMI NO₂ retrievals; Vasilkov et al., 2021); satellite data processing levels indicate the degree of data processing of satellite data, where L2 indicates that a geophysical parameter has been derived.

4.2 Thermal infrared (TIR) data

Trace gas measurements in the TIR spectrum often target species with relatively weak absorption, and low instrument noise is thus a key requirement for useful retrievals (Fu et al., 2019). High spectral resolution is also pivotal for resolving relevant spectral features and distinguishing them from atmospheric or surface interferences (Clarisse et al., 2011). In addition to providing spectral coverage encompassing the chemical species of interest, future TIR sounders aiming to constrain atmospheric composition therefore need to prioritize these instrumental characteristics.

New retrieval developments employing ML and species detection based on a hyperspectral range index (Walker et al., 2011) have led to increased sensitivity and expanded the suite of species that can be measured in the TIR range (e.g., acetic acid, acetone, ethane, ethene, ethyne, formic acid, glycoaldehyde, hydrogen cyanide, isoprene, methanol, peroxyacetyl nitrate (PAN)) (e.g., Wells et al., 2020; Franco et al., 2022, 2024; Brewer et al., 2024). Since TIR observations are based on the thermal contrast between the Earth surface and the absorber, uncertainties increase when that temperature difference is small. For the same reason, profile shape uncertainties can be an important error source in TIR retrievals. Validation data are often limited for species targeted in this spectral range, and future space-based measurements would benefit from an expanded portfolio of surface and airborne data to validate retrievals and refine profile shape assumptions. Observational errors can also arise from uncertainties in the absorption cross section and surface emissivity datasets employed in retrievals; advancing the fidelity of these resources would therefore benefit the reliability of TIR trace gas products.

4.3 Short-wave infrared (SWIR) data

Measurements of trace gases in the SWIR rely on reflected solar radiance and can provide a true total column observation through the atmosphere under cloud-free conditions and when the atmospheric pathlength is characterized. Here we assume the SWIR range includes wavelengths from approximately 0.7 to 3 μm. This range is especially useful for measuring total columns (surface to top of atmosphere) of CO₂ (Greenhouse gases Observing SATellite, GOSAT, and Orbiting Carbon Observatory-2, OCO-2), CH₄ (GOSAT and Tropospheric Monitoring Instrument, TROPOMI) and CO (MOPITT, GOSAT-2, and TROPOMI). In many cases, the O₂ A-band around 0.76 μm is used as a reference observation of the atmospheric path and cloud detection due to the constant and well-known concentration of molecular oxygen. SWIR observations of total column CO and CH₄ are especially important for quantifying surface source emissions that have a significant contribution to OH sinks (Sect. 5.2).

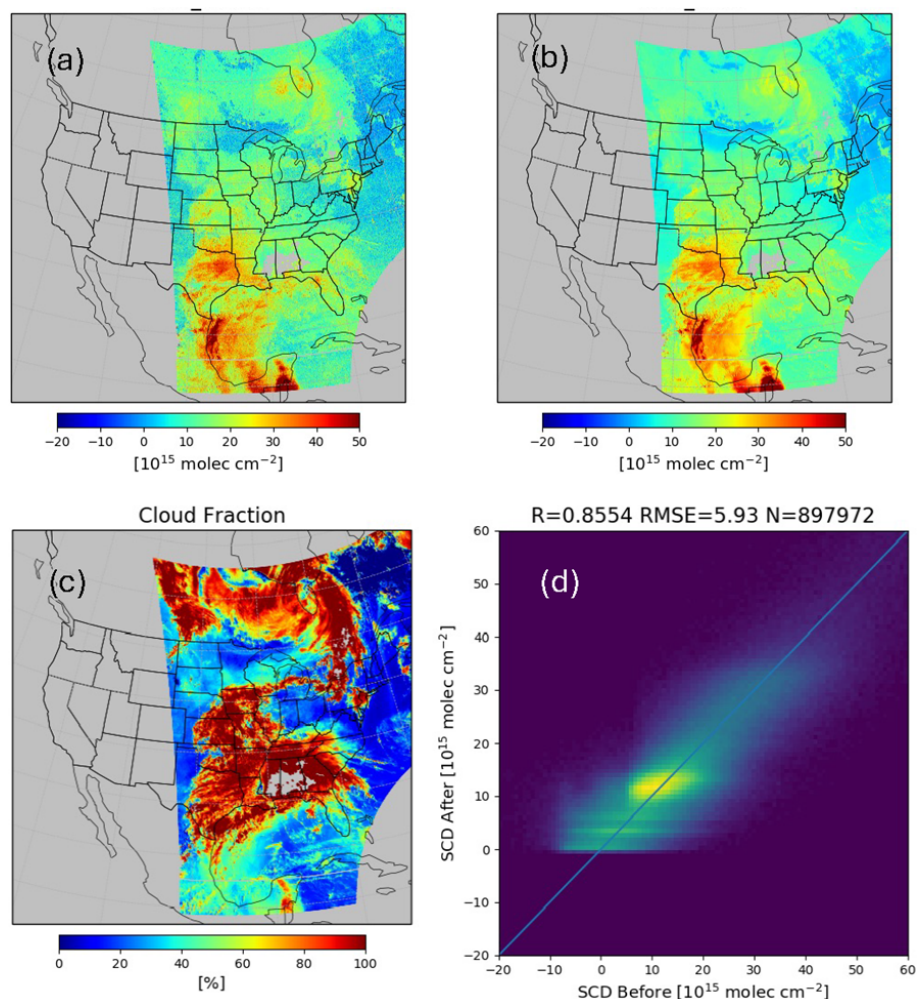


Figure 2. Preliminary results of noise reduction applied to HCHO-retrieved SCD from version 3 TEMPO data on 5 May 2024 scan 5: (a) original data before noise reduction used as training, (b) after noise reduction was applied, (c) effective cloud fraction supplied in the L2 data, and (d) the scatter diagram showing overall agreement between noise-reduced and original SCDs relative to the 1 : 1 line with fit statistics. R is the correlation, RMSE is the root-mean-square error, and N is the number of pixels used in the comparison.

4.4 Enhancing vertical resolution

As is the case for suborbital observation needs (Sect. 3), vertically resolved data of OH drivers are a priority for future satellite instruments to improve retrieval algorithms, especially for a priori profiles. In addition, they will advance the ability of the satellite-based approaches (Sect. 2) to constrain tropospheric OH at a finer vertical resolution than a tropospheric VCD measurement can provide. Such data are also useful for validating CTMs, developing process-based diagnostics (Sect. 5.3), and interpreting the causes of trends and variations in inferred OH throughout the tropospheric column (e.g., Baublitz et al., 2023). For example, the efficiency and quality of the ML approaches (Sect. 2.2) would benefit from vertically resolved input variables.

We recommend that studies be done to maximize the vertical information of the observations of OH drivers, which

may be achieved through a number of approaches, including developing satellite instrument technology (e.g., multi-angle and/or multispectral instruments) and investing in the development of multispectral data products that use radiances from one or more instruments (e.g., CH_4 ; Zhang et al., 2018). Multispectral data products were developed for O_3 (Tropospheric Emission Spectrometer, TES/OMI; Fu et al., 2013; Colombi et al., 2021; Atmospheric Infrared Sounder (AIRS)/OMI, Fu et al., 2018; TROPOMI/Cross-track Infrared Sounder (CrIS), Mettig et al., 2022) and CO (MOPITT NIR/TIR, Worden et al., 2010; Gaubert et al., 2017; CrIS/TROPOMI, Fu et al., 2016); development of these and other products is occurring through the NASA Tropospheric Ozone and Precursors from Earth System Sounding (TROPESS) project.

In addition to gaining vertical information through the development of satellite instrument technology and multi-

spectral algorithms, one may combine multiple data products of an OH driver from instruments that observe different tropospheric vertical layers. As an example, Oman et al. (2013) constrained the response of O₃ to ENSO using vertically resolved data from the NASA TES and MLS, both on the NASA Aura satellite. In this analysis, MLS and TES data were treated separately and not merged (TES operations ended in 2018 and MLS operations are predicted to end in mid-2026, with no follow-on instruments currently planned).

5 Recommendations for satellite needs

In this section, we discuss considerations and make recommendations for an observing strategy that would advance our ability to quantify the concentrations, trends, and variations in tropospheric OH using atmospheric remote-sensing observations of its drivers.

5.1 Optimizing an observing strategy through trade-space analyses

As discussed in Sect. 2, there are a number of approaches in the literature to indirectly constrain OH with satellite proxies for OH drivers. Each of the approaches applies to specific regions and/or photochemical environments or was devised for specific applications, and each of them has its strengths and limitations that should be considered. Therefore, some combination of these approaches, as well as the development of new approaches for specific photochemical environments, may be required to maximize spatio-temporal coverage of tropospheric OH around the globe for individual applications (as discussed below).

The design process of a comprehensive, multi-satellite observing strategy would benefit from trade-space analyses (i.e., assessing the impact of changing one or more variables while simultaneously changing one or more different variables in the opposite direction). For example, such analyses of spatial, temporal, and spectral resolutions for specific instruments would inform design decision making. The design process would also benefit from cost–benefit analyses. For instance, what observations are achievable using more cost-effective suborbital instruments (Sect. 3) than satellite instruments?

5.1.1 Continuity

The fidelity of a sustained, long-term record of inferred tropospheric OH from satellite-based approaches (Sect. 2) depends on the continuity of satellite proxies for OH drivers. Therefore, a priority for a robust observing strategy is for the continuity of satellite instruments with similar or enhanced capabilities, as well as inter-agency coordination, to maintain self-consistent and trend-quality level 1B (L1B) radiances for the generation of multi-instrument data records. L1B data, which are used in scientific algorithms to derive geophysical

parameters, have been calibrated and geolocated. The transition from one instrument to its successor instrument ideally would include a period of observational overlap (e.g., 2–5 years) so that self-consistent, multi-instrument data records may be achieved. As an example, near-daily global coverage of the NO₂ VCD is currently observed by OMI (since 2004) and its successor, TROPOMI (since 2017) as well as OMPS on Suomi NPP (since 2011), NOAA-20 (since 2017), and NOAA-21 (since 2022). All five instruments have overpass times near 13:30 local time. However, the OMPS instruments are missing visible wavelengths (e.g., < 380 nm for Suomi NPP) as compared to OMI or TROPOMI and have lower spectral resolutions (~ 1 nm), which result in lower quality NO₂ data products than those from either OMI or TROPOMI. The planned successor to TROPOMI, Sentinel-5, will have similar capabilities as TROPOMI but will have an Equator overpass time near 09:30 local time. Therefore, there will be discontinuities in the long-term, record given that NO_x emissions, NO_x partitioning between NO and NO₂, and the vertical distribution of NO₂ within the VCD vary throughout the day. Depending on TROPOMI's lifespan, potential successor instruments with overpass times near 13:30 local time, are the NOAA series of OMPS instruments being planned for the Joint Polar Satellite System (JPSS)-3 and JPSS-4 satellites. JPSS-3 and JPSS-4 OMPS will observe additional visible wavelengths relative to those observed by earlier OMPS instruments, but they will not have comparable instrument capabilities (e.g., SNR, spectral coverage) to OMI or TROPOMI. Therefore, the data quality of a self-consistent, multi-instrument data record for NO₂ at an overpass time near 13:30 local time from OMI, TROPOMI, and OMPS will be determined by the instrument parameters (e.g., spectral resolution, wavelength range) of the least capable OMPS instrument.

5.1.2 Accuracy/precision and signal-to-noise ratio (SNR)

As discussed in the scientific papers describing the satellite-based approaches (Sect. 2), additional improvements to accuracy/precision and SNR of current observing capabilities are required to advance our ability to constrain OH drivers. For instance, the design of new instruments should emphasize capabilities for in-flight instrument calibration and monitoring, such as solar diffusers and internal light sources. During the manufacturing and integration phase, adequate time and resources should be allocated to allow for thorough pre-flight calibration and characterization. As another example, technology improvements, such as better thermal design and detector cooling, can lead to reduced noise and enhancements in SNR.

5.1.3 Spatio-temporal coverage

In addition to enhanced instrument capabilities, improvements to current spatio-temporal coverage of satellite proxies for OH drivers could be achieved with additional satellites. For instance, the datasets used in the satellite-based approaches (Sect. 2) are from satellites in low-Earth orbit (LEO), so most global locations where OH levels are highest (i.e., tropics and subtropics) are observed daily or every few days. Therefore, a suite of instruments could be placed on multiple satellites, providing observations at various times throughout daylight hours. This strategy would also have the benefit of providing more opportunities to observe non-cloudy conditions (Sect. 5.4), especially if the instruments have finer spatial resolutions than current satellite instruments to allow more observations between clouds.

Instruments on geostationary (GEO) satellites provide an opportunity to assess the value of hourly data to constrain OH. For instance, several new UV/Vis instruments such as the Korean Geostationary Environment Monitoring Spectrometer (GEMS; Kim et al., 2020) and NASA's Tropospheric Emissions: Monitoring of Pollution (TEMPO; Zoogman et al., 2017) observe VCDs of NO₂, HCHO, and ozone over East Asia and North America, respectively. ESA's Sentinel-4 GEO satellite with the Ultraviolet–Visible–Near-Infrared (UVN) instrument will make similar observations over Europe and northern Africa (Courrèges-Lacoste et al., 2017). The spatial coverages of these instruments are primarily limited to the midlatitudes and subtropics of the Northern Hemisphere. For hyperspectral TIR instruments on GEO satellites, the Chinese Fengyun Geostationary Interferometric Infrared Sounder already exists (GIIRS; Yang et al., 2017); similar instruments are tentatively planned by space agencies in Japan (Okamoto et al., 2020), Europe (Holmlund et al., 2021), and the US (Lindsey et al., 2024). The spatial coverages of these instruments are more comprehensive, covering most latitudes (i.e., below $\sim 60^\circ$) of both hemispheres, as compared to the coverages of the UV/Vis instruments. We recommend coordinated deployment of UV, Vis, NIR, and TIR instruments on GEO satellites to provide co-located observations of satellite proxies for key OH drivers (Sect. 5.2).

5.1.4 Case studies

Ultimately, the design of a space-based observing strategy to constrain tropospheric OH depends on the specific science need or needs to be addressed. For instance, is there a need to constrain OH concentrations, trends, and/or variations, which could require different designs? For what regions and seasons would this be necessary? For illustrative purposes, we provide design considerations for two different potential applications.

In case 1, we wish to constrain the spatio-temporal variations in methane's loss by reaction with OH, which occurs primarily in the lower troposphere of the tropics and

subtropics. Given methane's long lifetime, one might think that the spatial and temporal resolutions of inferred tropospheric OH could be low (e.g., several degrees latitude by longitude, seasonal in photochemically homogeneous environments). However, one issue that requires consideration is that the number of pixels with sufficiently low cloud contamination would decrease as the horizontal resolution of observations degrades in cloud-prone areas (e.g., in much of the tropics and subtropics; Sect. 5.4). Therefore, an observational strategy for this case would require finer spatial resolutions for measurements of OH drivers than are required for the spatial resolution of inferred OH. In addition, much of the loss of CH₄ occurs over relatively remote ocean environments, so improving the SNR of OH drivers is particularly important, especially since the signals for some important species (e.g., NO₂, HCHO) are on par with noise for current instruments (e.g., Zhu et al., 2020). Potentially, new satellite missions can be designed to take advantage of sun glint, where specular reflection over water bodies can enhance the SNR. Over land in the tropics and subtropics, observations of additional VOC constituents, especially isoprene (e.g., Fu et al., 2019; Shutter et al., 2024), are desired to constrain tropospheric OH (Sect. 5.2). Given the strong spatial gradients of OH drivers, spatial resolution should also be prioritized over land as well as over ocean affected by offshore flow. Satellite observations of dynamical variables (e.g., meteorological, climate indices) may also be useful for this case as they can influence OH (Sect. 5.3). For this case, design objectives may be met by multiple instruments with different design objectives (e.g., fine spatial resolution for over land vs. improved SNR for over ocean) or instruments with multiple sampling modes.

In case 2, we wish to constrain tropospheric OH in polluted urban areas. This environment is typically heterogeneous and complicated, requiring fine spatio-temporal resolutions (e.g., kilometers, hourly, or daily) to capture spatial gradients in OH drivers. Additional VOC constituents are highly desired as the urban environment has thousands of VOCs; therefore, observations of VOCs that represent classes of VOCs (e.g., isoprene, fast-reacting alkenes, alkanes, aromatics) are desired and will require multispectral satellite observations (i.e., TIR, NIR, UV/Vis) for the development of these data products (Sect. 5.2). Depending on the timescale of interest, photochemical variables may be sufficient to constrain OH, which can be assumed to be in a photostationary state given its short lifetime (< 1 s). Nevertheless, meteorological variables that influence OH (e.g., temperature, convection) may be desired for studying the photochemical evolution of the urban plume by hour, for instance. For this case, design objectives may be met by observations of UV/Vis (e.g., NO₂, HCHO) and TIR (VOCs) wavelengths. Instruments on GEO satellites would provide fine temporal resolution, as opposed to the ones on LEO satellites.

5.2 Assessment of current satellite data to constrain chemical processes that influence OH

We recommend studies to determine the optimal combination of input variables/satellite observables for the satellite-based approaches (Sect. 2) to maximize the interpretability (in a process-based sense) of the causes of trends and variations in tropospheric OH. ML approaches (Sect. 2.2) are able to reproduce OH quite well from the training datasets using various satellite observations; however, the utility of the approaches for attribution studies is determined by the choice of input variables. That is, the suite of input variables must adequately represent the chemical processes of interest that determine OH. For instance, attribution of OH variations and trends to the influence of CH₄ is not possible in an ML study if CH₄ concentration is not included as an input variable (e.g., Anderson et al., 2024). In addition, ML techniques may not reproduce OH in a process-based sense if the input variables to the ML model are not independent of one another, which is an inherent characteristic of atmospheric chemistry problems (e.g., O₃ and NO₂, temperature-dependence of most species' lifetimes, co-emitted species such as CH₄ and VOCs, and tropospheric species impacted by the same meteorology). Therefore, the causes of trends and variations in tropospheric OH inferred from ML approaches should be carefully interpreted. This is also true for process-based approaches (Sect. 2.1). For example, Pimlott et al. (2022) concluded that their approach “suggests that O₃ and CO were the key drivers of variability in the production and loss of OH” for their study period. However, their approach uses input variables observed from a single satellite instrument that does not collect data on NO₂, which is responsible for a substantial portion of the variations in tropospheric OH (e.g., Zhu et al., 2022a; Anderson et al., 2023; Baublitz et al., 2023).

5.2.1 A thought experiment

A key question is as follows: what chemical processes are not adequately represented in the current suite of satellite observations that are used in the satellite-based approaches (Sect. 2) to indirectly constrain local and regional concentrations of and trends and variations in tropospheric OH? To answer this question, it is instructive to assess our current ability to constrain individual sources and sinks for tropospheric OH (Table 1). For this thought experiment, we use the annual tropospheric OH budget presented in Table 1 of Lelieveld et al. (2016), noting that the relative importance of the sources and sinks for various regions and seasons can vary substantially from the annual global distribution. In addition, we acknowledge that OH budgets can vary significantly from one CTM to another (Sect. 1). As we will show, the results of our thought experiment, as summarized in Table 1, reveal that the ideal observing strategy for the indirect constraint of the

chemical processes that control OH requires co-located observations of the UV/Vis, NIR, and TIR wavelengths.

The first and second sources (i.e., O(¹D) + H₂O_(v)) and NO + HO₂; Table 1) account for 63 % of the total source in Lelieveld et al. (2016) and are represented relatively well with current satellite datasets. For the second source, satellite observations of NO₂ can be used to estimate NO using a CTM; this conversion is a function of a number of variables such as ozone concentrations and photolysis (e.g., de Foy et al., 2015). Lightning is an important source of NO_x in the middle and upper troposphere (Allen et al., 2021), which modulates OH there (e.g., Fiore et al., 2006). Satellite observations of lightning flash counts (e.g., over the Americas from the Geostationary Lightning Mapper, GLM, aboard the NOAA Geostationary Operational Environmental Satellite-16, GOES-16 and over Europe, Africa, and the Middle East from EUMETSAT Meteosat Third Generation – Imager 1 (MTG-I1) Lightning Imager) may be useful to constrain the vertical distribution of NO₂ using ML or a CTM to relate flash counts to NO₂ concentrations. While we do have estimates of tropospheric O₃ VCDs, representation of the third source (i.e., O₃ + HO₂) would benefit from an accurate separation of the stratospheric and tropospheric portions of the total O₃ VCD (e.g., Ziemke et al., 2006; Orfanoz-Cheuquelaf et al., 2024) in addition to information on the vertical distribution of tropospheric O₃ (Sect. 4.4). We are unaware of satellite proxies for tropospheric HO₂, which occurs in both the second and third sources and has a lifetime that is slightly longer than the lifetime of OH. Field campaign data from remote ocean regions indicate that HO₂ variation does modulate the rates of the second and third OH source reactions (Fig. 3a, b), but whether possible proxies for HO₂ (Fig. 3c) are robust in more complex photochemical regimes is a question that requires further research. A proxy for HO₂ may not be necessary for the ML approaches (Sect. 2.2), given that OH and HO₂ (i.e., HO_x family) share many of the same chemical drivers. There are satellite observations from the Atmospheric Chemistry Experiment (ACE) for H₂O₂, the fourth source (i.e., H₂O₂ + *hν*), in the middle to upper troposphere, although this source primarily occurs in the lower troposphere (e.g., Spivakovsky et al., 2000); these observations are sparse, so multi-year averages are required to obtain seasonal zonally averaged distributions (Allen et al., 2013). For satellite proxies that are not yet well supported by current satellite instruments, such as H₂O₂ and many VOCs (discussed below), simulated distributions may be used in the satellite-based approaches (Sect. 2), especially the distributions that have been well validated with suborbital datasets (Sect. 3). Similar to the argument for HO₂, OH and H₂O₂ share many of the same chemical drivers, so we may not need a satellite proxy for H₂O₂ in ML approaches (Sect. 2.2). We will discuss the fifth source, which involves VOCs, below. We added the source of OH from the photolysis of nitrous acid (HONO; HONO + *hν*) to Table 1, although it is not included in Table 1 of Lelieveld et al. (2016). The over-

all importance of this source on a global scale is uncertain (e.g., Wu et al., 2022; Ha et al., 2023; Zhang et al., 2023), but is thought to be an important source in urban and agricultural environments. Satellite observations of HONO (e.g., TROPOMI, IASI) are primarily limited to intense wildfire plumes (e.g., Theys et al., 2020; Fredrickson et al., 2023; Franco et al., 2024).

The sinks that have straightforward satellite proxies are the second and third sinks (i.e., OH + CH₄; OH + CO; e.g., Gaubert et al., 2017; Nguyen et al., 2020; Jacob et al., 2022; Worden et al., 2022), which together represent 51 % of the total tropospheric OH sink from Lelieveld et al. (2016). CH₄ is relatively well mixed throughout the troposphere; therefore we could assume a tropospheric distribution informed by in situ observations. For the first sink (i.e., OH + HO_y, where HO_y = H₂, O₃, H₂O₂, radical–radical reaction), there are satellite data for tropospheric O₃ and limited data for H₂O₂, as discussed above. In addition, H₂ is relatively well mixed throughout the troposphere; therefore we could assume a tropospheric distribution informed by in situ observations. We will discuss the fourth and fifth sinks, which involve VOCs, below.

Overall, we currently have reasonable satellite proxies to constrain up to 77 % of the global source and up to 51 % of the sink (Table 1) in the annual tropospheric OH budget of Lelieveld et al. (2016). With the development of new VOC data products (e.g., De Longueville et al., 2021; Franco et al., 2022; Brewer et al., 2024; Wells et al., 2024), which may require new technology development, we may well be able to represent a larger portion of the total sources (up to 90 %) and sinks (up to 80 %). To do this would require that we leverage numerous spectral regions (UV/Vis, NIR, TIR) to maximize our ability to represent key source and sink reactions for tropospheric OH with satellite proxy data. An important caveat is that the satellite data products in Table 1 are not of equal quality and require additional research and/or technological development to improve their utility for indirectly constraining OH. A quantitative assessment of each data product's quality should be performed using independent suborbital observations when available (Sect. 3). Such an assessment would benefit from inter-comparable uncertainty characterization between the different data products; however, this is not currently done. Therefore, we recommend the use of a common set of reporting standards to be applied for uncertainty characterization of satellite data products (e.g., von Clarmann et al., 2020), which will allow for a more inter-comparable assessment of the utility of each satellite data product for indirectly constraining tropospheric OH.

We recommend that additional research be a priority to explore the potential of these observable VOCs to constrain tropospheric OH. In Table 1, there is one source (i.e., OVOCs, ROOH + *hν*; 13 % of total source) and two sinks (i.e., OH + other C₁ VOC and OH + C₂₊ VOC; 29 % of total sink) that would benefit from the development of space-based observations of VOCs, including oxygenated VOCs (OVOCs). Several VOCs are currently observable from space, including in the UV/Vis (HCHO, glyoxal) and TIR (acetic acid, acetone, ethane, ethene, ethyne, formic acid, glycoaldehyde, hydrogen cyanide, isoprene, methanol, and peroxyacetyl nitrate (PAN); e.g., Franco et al., 2018, 2019, 2020, 2022; Wells et al., 2024). Given that it is not feasible to observe the hundreds of individual VOCs in the troposphere and potentially thousands in the polluted urban atmosphere, we recommend that observable VOCs be identified that could represent (1) important classes of compounds (e.g., alkanes, alkenes, aromatics) that contribute to total OH reactivity and (2) markers of various VOC emission sources (e.g., biogenic, anthropogenic, pyrogenic), which have characteristic VOC mixes. Additional constraints of these sources and sinks may benefit from other satellite proxies such as CO (e.g., Baublitz et al., 2024) that covary with VOCs under certain conditions.

We have satellite proxies for the dependence of reactions on temperature and sunlight. The temperature dependence of chemical reactions is represented by satellite observations of temperature (TIR), and continuity is assured with CrIS, for example. The stratospheric O₃ VCD (UV/Vis) may serve as a proxy for sunlight (*hν*), where photochemically important wavelengths are modulated by the thickness of the stratospheric O₃ layer. Continuity is assured with several current and upcoming UV/Vis instruments (e.g., TROPOMI, OMPS). Aerosols also modulate the amount of sunlight in the troposphere, affecting tropospheric oxidants (e.g., Martin et al., 2003; Mok et al., 2016; Madronich et al., 2024). There are several potential proxies, such as total and absorption aerosol optical depth and single scattering albedo, that are observed by a number of satellite instruments (e.g., on ESA's Earth Cloud Aerosol and Radiation Explorer (EarthCARE) satellite and NASA's Plankton, Aerosol, Cloud, ocean Ecosystem (PACE) satellite). Additional research is required to develop a combined satellite proxy for spectral solar planar and scalar (i.e., actinic flux) irradiances (e.g., Vasilkov et al., 2022).

In summary, the results of our thought experiment, as summarized in Table 1, reveal that the ideal observing strategy for the indirect constraint of the chemical processes that control OH would require co-located observations of multiple wavelengths in the UV/Vis, NIR, and TIR spectra, which could be integrated into a broader, comprehensive observation strategy for tropospheric chemistry and climate (Millet et al., 2024).

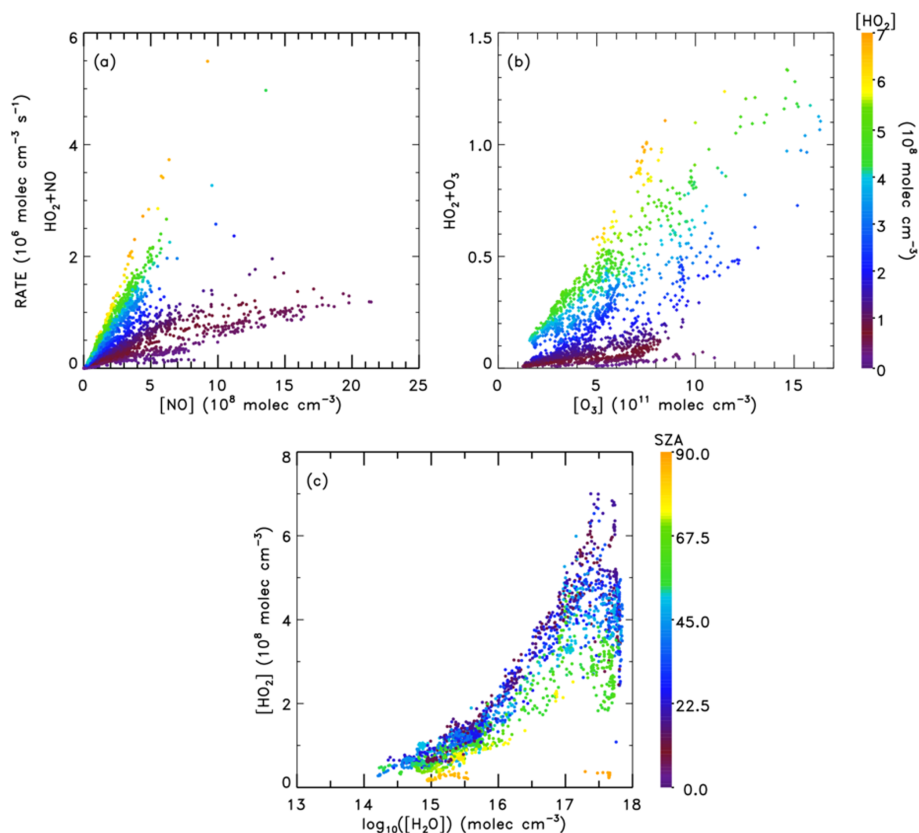


Figure 3. Data shown are for tropical over-ocean samples from the ATom-1 field campaign. Box-modeled rates are for the $\text{HO}_2 + \text{NO}$ reaction plotted against the NO concentration (panel **a**) and the $\text{HO}_2 + \text{O}_3$ reaction vs. the O_3 concentration (**b**), both colored by concentration of HO_2 . Panel (**c**) shows the HO_2 concentration plotted against \log_{10} of water vapor concentration and colored by solar zenith angle.

5.2.2 Looking backward

This paper is primarily focused on looking forward – i.e., what is required/desired to improve our ability to constrain tropospheric OH with satellite observations? However, a natural question is how far back in time can we adequately constrain tropospheric OH. Addressing this question and comparing estimates of past OH abundance would increase confidence in our ability to project OH into the future. Answering this question depends on the level of explanatory power that is desired, as discussed above. For instance, one of the first satellite proxies listed in Table 1 is the stratospheric O_3 VCD. As shown by Rohrer and Berresheim (2006), tropospheric OH correlates linearly with solar UV irradiance, which is modulated by stratospheric O_3 . Therefore, one could partially constrain tropospheric OH back to 1979, when satellite measurements of stratospheric O_3 became routine (e.g., Stolarski et al., 1986; Weber et al., 2018). As another example, discussion surrounding the continuity of the NO_2 VCD (Sect. 5.1.1) started with OMI (launched in 2004), although such observations actually began in 1996 with ESA’s Global Ozone Monitoring Experiment (GOME, 1995–2011; Burrows et al., 1999), which was followed by SCanning Imaging Absorption SpectroMeter for Atmospheric CHartography

(SCIAMACHY, 2002–2012), GOME-2 instruments on the Meteorological Operational (METOP) satellites (METOP-A, 2006–2021; B, since 2012; and C, since 2018). All these ESA instruments were/are in morning orbits. Therefore, depending on one’s tolerance for data that have relatively coarser resolution, poorer SNR, or a longer time period to obtain global coverage, partial constraint of tropospheric OH with satellite data of NO_2 from a morning orbit could extend back to 1996 (e.g., Boersma et al., 2018). Continuity of the morning orbit is planned with Sentinel-5 (anticipated launch in 2025).

Continuity is assured for most satellite proxies (Table 1), and we currently have the ability to adequately or partially constrain most of the sources and sinks of tropospheric OH (Table 1) with the satellite-based approaches (Sect. 2) from the late 1990s/early 2000s to the present. As discussed in Sect. 5.2.1, the primary exception is related to VOCs (e.g., isoprene, which only became available in 2012).

5.3 Influence of tropospheric dynamics on OH

We recommend additional research to determine the full potential of using satellite-observable meteorology (e.g., tem-

Table 1. Satellite proxies for tropospheric source and sink fluxes of OH.

	Sources ^a	% of total sources ^a	Satellite proxies (wavelengths)	Satellite proxy limitations and potential improvements	Short-term data continuity? ^h
1	O(¹ D) + H ₂ O _(v)	33 %	Stratospheric O ₃ VCD (UV/Vis) for O(¹ D) or assume steady-state: [O(¹ D)] = $j[\text{O}_3]/(k_a[\text{H}_2\text{O}] + k_b[\text{O}_2] + k_c[\text{N}_2])$ using tropospheric O ₃ VCD. Water vapor (IR).	No major limitations.	Tropospheric & stratospheric O ₃ VCD (yes) – e.g., TROPOMI. Water vapor (yes) – e.g., CrIS.
2	NO + HO ₂	30 %	NO ₂ VCD (UV/Vis). Research is needed to identify a proxy for HO ₂ .	Needs improved SNR where NO ₂ VCDs are low (Buscela et al., 2013). Lightning flash counts may provide information on the vertical distribution (e.g., NOAA GeoXO Lightning Mapper (LMX) or MTG-II Lightning Imager), but data are typically limited to certain regions and time periods.	NO ₂ VCD (yes) – e.g., TROPOMI.
3	O ₃ + HO ₂	14 %	Tropospheric O ₃ VCD (UV/Vis). Multispectral products (Sect. 4.4). Research is needed to identify a proxy for HO ₂ .	Needs accurate stratosphere–troposphere separation of total column O ₃ VCD (Ziemke et al., 2006). Research is needed to determine the potential of multispectral products (e.g., TROPOMI/CrIS), which may provide information on vertical distribution.	Tropospheric O ₃ VCD (yes) – e.g., TROPOMI, OMPS.
4	H ₂ O ₂ + <i>hν</i>	10 %	H ₂ O ₂ (IR).	Instrument and/or retrieval algorithm development required, including to obtain observations in the lower troposphere. ACE data of H ₂ O ₂ are sparse; zonal averages of multiple years are required to obtain near-global coverage in the middle to upper troposphere (Allen et al., 2013).	H ₂ O ₂ (no).
5	OVOCs ^b , ROOH ^c + <i>hν</i>	13 %	HCHO, glyoxal (UV/Vis). Numerous potential VOCs (TIR).	Needs improved SNR where HCHO VCDs are low. VOC instrument (TIR) and retrieval algorithm development required. Research required to determine the suitability of these VOCs for constraining this source.	HCHO, glyoxal VCD (yes) – e.g., TROPOMI. Numerous VOCs (yes for some but not others) – e.g., CrIS.
	HONO ^d + <i>hν</i>	–	HONO (UV/Vis, TIR).	Observations primarily for intense wildfire plumes that reach altitude. Data are very noisy. Instrument and/or retrieval algorithm development required.	HONO (yes) – e.g., TROPOMI.
	Sinks ^a	% of total sinks ^a			
1	OH + HO _y ^e	18 %	Tropospheric O ₃ VCD (UV/Vis). H ₂ O ₂ (IR). Assume constant distribution of H ₂ .	See source no. 4 above.	See source no. 3 above.
2	OH + CH ₄	12 %	CH ₄ (IR).	Needs (1) better SNR to detect variations in high background concentration and (2) improved sensitivity to near-surface.	CH ₄ (partially yes) – e.g., TROPOMI (over land, glint over ocean).
3	OH + CO	39 %	CO (IR). Multispectral products (Sect. 4.4).	Needs improved near-surface sensitivity. Research is needed to determine the potential of multispectral products (e.g., TROPOMI/CrIS), which may provide information on vertical distribution.	CO VCD (yes) – TROPOMI, CrIS.

Table 1. Continued.

	Sinks ^a	% of total sinks ^a	Satellite proxies (wavelengths)	Satellite proxy limitations and potential improvements	Short-term data continuity? ^h
4	OH + other C ₁ VOC ^f	15 %	Methanol.	VOC instrument (TIR) and retrieval algorithm development required. Research required to determine the suitability of these VOCs for constraining this sink.	Numerous VOCs (yes for some but not others) – e.g., CrIS.
5	OH + C ₂₊ VOC ^g	14 %	Isoprene, PAN, etc.	VOC instrument (TIR) and retrieval algorithm development required. Research required to determine the suitability of these VOCs for constraining this sink.	Numerous VOCs (yes for some but not others) – e.g., CrIS.

^a Reproduced from Table 1 of Lelieveld et al. (2016), except neglecting two minor sinks. ^b OVOCs = oxygenated VOCs, such as acetone and acetaldehyde. ^c ROOH = organic peroxides, such as CH₃OOH. ^d HONO is not explicitly listed as a source of OH in Lelieveld et al. (2016), although it is an important source in some environments (Theys et al., 2020; Fredrickson et al., 2023). ^e HO_y = H₂, O₃, H₂O₂, radical–radical reaction. ^f VOC with one C atom (excluding methane), including methanol and C1-reaction products. ^g VOC with ≥ 2 C atoms, C₂₊-reaction products. ^h Defined in Sect. 5.1.1. We focus on current satellite instruments that provide near-complete spatial coverage of the troposphere, neglecting instruments that provide data of, for instance, the upper troposphere (e.g., MLS) and geostationary orbits. We do not list potential future instruments given that space agency priorities and budgets change and because of the history of satellites failing to reach orbit during launch.

perature) and climate variables in space-based approaches (Sect. 2) to better understand the dependence of tropospheric OH on dynamics. For instance, climate indices, such as the Multivariate El Niño–Southern Oscillation (ENSO) Index (MEI) and indicators of drought, correlate well with long-term (e.g., monthly, seasonal) variations in tropospheric constituents (e.g., CO, NO₂, O₃, isoprene, water vapor) that influence OH (e.g., Oman et al., 2011, 2013; Turner et al., 2018; Wells et al., 2020; Anderson et al., 2021, 2023; Shutter et al., 2024), as well as OH itself. Anderson et al. (2021) estimated that ~ 30 % of simulated variability in global OH in boreal winter is associated with ENSO alone. However, additional proxies may be required to fully capture the complex impact of meteorological variations on the tropospheric constituents that influence OH, including weather-dependent emissions (e.g., CH₄, isoprene, lightning NO_x) and anthropogenic activities (e.g., Shutter et al., 2024). For instance, Field et al. (2016) reported a nonlinear sensitivity of CO abundance to emissions from fires in Indonesia during El Niño-induced drought, when many fires escape or are intentionally set to clear land on areas with degraded peat. Building on this research, the locations of peat deposits convolved with satellite observations of fire-counts could be used to estimate pollution emitted from these fires. Alternatively, a total sum of pollutant emission estimates (e.g., Global Fire Emissions Database, GFED; van Wees et al., 2022) from Indonesia or the preceding dry season total rainfall in conjunction with the locations of degraded peat could be used to capture the impact of ENSO variations on tropospheric OH.

We recommend additional research to develop process-based diagnostics using satellite and suborbital observations to improve the representation of key atmospheric processes in CTMs that influence OH. Although many of the current knowledge gaps in CTMs (e.g., Table 2 of Fiore et al., 2024) require new laboratory studies, enhancements to current satellite capabilities as well as a more comprehensive suite of suborbital datasets (Sect. 3; e.g., Murray et al.,

2021) could be used to develop process-based diagnostics to better constrain the factors that influence OH drivers, such as constraining the response of O₃ to ENSO (Oman et al., 2013). Such diagnostics would inform development priorities of CTMs with the goals of improving the simulation of tropospheric OH and aiding interpretation of inferred OH from satellite-based approaches (Sect. 2).

5.4 Addressing the “clear-sky” bias of inferred OH from satellite-based approaches

The presence of clouds limits a satellite’s ability to observe tropospheric composition, especially in seasonally cloudy regions. King et al. (2013) used MODIS cloud fraction data to estimate that ~ 67 % of Earth’s surface is covered by clouds at any given time. Consequently, temporal averaging (e.g., over a season) of the data is often employed to achieve statistical significance, although the averages are biased toward clear-sky/low-cloud conditions. We found that these biases in simulated OH (from one CTM) relative to all-sky conditions are substantial (e.g., ±25 % in the tropospheric OH VCDs in the tropics and higher elsewhere; Fig. 4). One way to reduce these biases is to improve the spatial resolutions of satellite instruments, which would increase the chances of collecting observations between broken clouds and, thus, minimize the number of cloud-contaminated pixels (e.g., Frankenberg et al., 2024). There may be viable ways to infer some of the OH drivers over clouds (e.g., NO₂; Marais et al., 2018, 2021; CO, Landgraf et al., 2016), but further study is needed.

We recommend additional research to identify ways to gap fill OH inferred from the satellite-based approaches (Sect. 2). Potential ways are to (1) use simulated OH, which would benefit from model development priorities informed by satellite-constrained, process-based diagnostics (Sect. 5.3); (2) scale inferred OH by the ratio of observed OH in clear and cloudy conditions from tropospheric composition field campaigns (Sect. 3); and (3) use a photolysis model with satellite-observable cloud properties (e.g., cloud

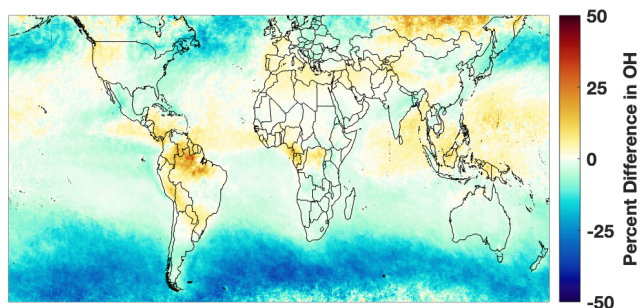


Figure 4. Simulated percent difference in tropospheric column OH between low-cloud (cloud fraction less than 30 %) and all-sky conditions. Red indicates higher OH abundance in near-clear-sky conditions; blue indicates lower OH abundance under such conditions. Simulated data are from the MERRA-2 GMI simulation and averaged over 2005–2019 for October.

optical depth, cloud top pressure, and cloud fraction) to scale inferred OH. The problem with the second way is that cloudy regions are typically avoided during field campaigns. Therefore, a priority of future field campaigns is to collect observations of OH and its drivers in both clear and cloudy environments. Finally, there is the potential to develop an ML model using simulated tropospheric OH with meteorological variables and climate indices (Sect. 5.3), which are observable from space, to understand the relationships of OH concentrations, trends, and variations in cloudy and mostly cloudy environments. All of these potential methods to gap fill inferred OH require further study.

6 Summary of recommendations

The satellite-based approaches (Sect. 2) offer the promise that we may achieve constraint of tropospheric OH on spatio-temporal scales that has been previously unobtainable. Given this promise, we make recommendations to facilitate the design of a robust observing strategy of satellite and suborbital observations of atmospheric constituents and dynamical variables to indirectly constrain concentrations of and trends and variations in tropospheric OH. The increased spatio-temporal information about tropospheric OH that could be gained by following these recommendations would advance our ability to determine the causes of trends and variations in the concentrations of atmospheric constituents that are influenced by OH. For instance, Turner et al. (2016) concluded that “the current surface observing system does not allow unambiguous attribution of the decadal trends in methane without robust constraints on OH variability, which currently rely purely on methyl chloroform data and its uncertain emissions estimate.”

We recommend the development of a comprehensive satellite-based approach to maximize the indirect constraint of tropospheric OH that may be composed of a combination of existing satellite-based approaches (Sect. 2) and new satellite-based approaches developed for specific photochemical environments. Studies are needed to determine the optimal combination of input variables/satellite observables for some of the satellite-based approaches to maximize the interpretability (in a process-based sense) of the causes of trends and variations in tropospheric OH. The design process would benefit from trade-space and cost–benefit analyses, including consideration of incorporating more cost-effective suborbital observations where feasible.

Priorities (in no particular order) for this observing strategy include the following:

- continuity of satellite instruments with similar or enhanced capabilities;
- improvements to accuracy/precision and SNR of current observing capabilities through instrument design;
- additional LEO and GEO satellites to obtain greater spatio-temporal coverage;
- development of a robust suborbital observing network composed of instruments at stationary sites (OH supersites) as well as ones deployed in field campaigns to collect suborbital observations of OH drivers, including in cloudy environments;
- a common set of reporting standards to be applied for uncertainty characterization of satellite data products, which will allow for a more robust assessment of the utility of each satellite data product for indirectly constraining tropospheric OH; and
- coordinated deployment of UV, Vis, NIR, and TIR instruments to provide co-located observations of satellite proxies for key OH drivers.

We also recommend additional investment in research (in no particular order) to

- maximize the vertical information of the observations of OH drivers (e.g., with multispectral data products);
- maintain self-consistent and trend-quality L1B radiances for the generation of multi-agency, multi-instrument data records;
- determine the full potential of using satellite-observable meteorology (e.g., temperature, clouds, aerosols) and climate variables in space-based approaches to better understand the dependence of OH concentrations, trends, and variations on dynamics;
- develop process-based diagnostics using satellite and suborbital observations to improve the representation of key atmospheric processes in CTMs that influence OH;

- explore the potential of observable VOCs to aid constraint of tropospheric OH; and
- address the clear-sky bias of inferred OH from satellite-based approaches.

Finally, we recommend additional investment in technology development that may lead to passive and/or active space-based instruments that directly observe OH.

Data availability. Figure 1 – data are available upon request from Daniel C. Anderson (daniel.c.anderson@nasa.gov).

Figure 2 – satellite retrievals for TEMPO HCHO V3 data can be found at https://doi.org/10.5067/IS-40e/TEMPO/HCHO_L2.003 and the radiance v3 data at https://doi.org/10.5067/IS-40e/TEMPO/HCHO_L2.003 (NASA/LARC/SD/ASDC, 2024a), and the irradiance v3 data at https://doi.org/10.5067/IS-40e/TEMPO/IRR_L1.003 (NASA/LARC/SD/ASDC, 2024b; González Abad et al., 2024).

Figure 3 – data from the ATom campaign are located at <https://doi.org/10.3334/ORNDAAC/1925> (Wofsy et al., 2021).

Figure 4 – output from the MERRA-2 GMI simulation is publicly available at <https://acd-ext.gsfc.nasa.gov/Projects/GEOSCCM/MERRA2GMI/> (NASA Goddard Space Flight Center, 2024).

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