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Maximum-likelihood fit to points originating from different Poisson distributions

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1. Introduction

In the determination of absolute alpha-particle energies in a magnetic spectrograph the energy is obtained by extrapolating the high energy edge of the alpha line to zero intensity (see, for example, ref. [l]) « The method consists of fitting to the experimental points a theoretical intensity distribution I_x proportional to $(x - a)^{3/2}$ where x is the position **on the photographic plate and a is the intercept. The procedure used** until now [2] has been to take I_x, the measured number of alpha tracks per unit interval, subtract from it the average background 1₀ and transform the resultant quantity $I_x - I_0$ to the power $2/3$. Then, using appropriate $\frac{1}{2}$ $\frac{1}{2}$ **weights, a straight line of the form x = a + b(l - I) is fitted by the method of least-squares. This proccdu.e has the advantage that it leads to** $\frac{1}{2}$ $\frac{1}{2}$ $\frac{1}{2}$ $\frac{1}{2}$ $\frac{2}{2}$ simple analytic expressions $\begin{bmatrix} 2 \end{bmatrix}$ for the intercept α and its variance $\begin{bmatrix} 0 \\ 0 \end{bmatrix}$

The total number of measured tracks I follows a Poisson distribution 2 *x* and hence its variance is given by $\begin{array}{cc} 0 & -1 \\ x & x \end{array}$, when setting error limits on this number, one commonly writes the best value as $\iota_x \pm \sigma_x$. However, this does not reflect the asymmetry of the Poisson distribution for which **this does not reflect the asymmetry of the Poisson distribution for which one would expect asymmetric error limits. For large values of I , this** asymmetry becomes quite small and can then be neglected, but in the
region where we are actually performing the fit (i.e. at the high energy **region v/here we are actually performing the fit (i.e . at the high energy edge of the alpha line), I approaches zero.**

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It would therefore seem that a more accurate extrapolation of the alpha line to zero intensity could be obtained by considering the Poissonian nature of the process and the fact that the errors about the measured numbers of alpha tracks will be asymmetric. The technique of least-squares is unable to treat this case of asymmetric errors and so a procedure has been adopted here based on the principle of maximum likelihood.

2. Determination of asymmetric errors

A treatment of the problem of determining confidence limits on rare events distributed according to the law of Poisson is given by van der Waerden [3] and we only quote here his results.

If k is the observed number of events of a process following a Poisson distribution law, then the confidence limits on the value k are given by

$$
\lambda_{\pm} = k + \frac{1}{2} g^2 \pm g \sqrt{k + \frac{1}{4} g^2} \tag{1}
$$

where λ_+ and λ_- are the upper and lower confidence limits, respectively. **The factor g determines the degree of confidence and is taken from a normal distribution. Thus, for example, g = 1 corresponds to confidence limits of about 68% . If the distribution were normal this would agree with the usual standard deviation.**

For small values of k the limits of (1) differ considerably from those obtained by simply taking $\sigma = \sqrt{k}$. A few values are given in the table **below for comparison. The value of g is taken as unity.**

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line shape to points with asymmetric errors. The question now arises as to what is the best fit of our theoretical $\frac{1}{\sqrt{2}}\frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right) ^{2}$

Table 1. Error limits for a small number of counts (from (1), for g = 1)

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3» Principle of the maximum likelihood

The method used in what follows is based on the principle of maximum likelihood. The measured points are designated by (x. , y.), where **x**₁ is the position of the centre of the measured interval on the **photographic plate, and y. is the number of tracks counted (including background) in this interval. We assume that there is no error in the x.** and that the error distribution in the y_i is Poissonian and thus given by

$$
p(y_i|\lambda_i) = \frac{\lambda_i^{y_i}}{y_i!} \cdot e^{-\lambda_i} \qquad (2)
$$

X. is the (unknown) true value of which y. is a measure, i ' i

If λ _: were known, we could calculate, using (2), the probability **of observing a certain number of tracks y.. However, we have only the** measured quantity y_i , and wish to use it to estimate λ_i by means of the inverse probability $p(\lambda_i | y_i)$, for the distribution of the unknown true values λ_i .

i *•* **This can be done by applying the method of Bayes £4jj which gives**

$$
p(\lambda_{i}|y_{i}) = \frac{p(\lambda_{i}) \cdot p(y_{i}|\lambda_{i})}{\int_{0}^{\infty} p(\lambda_{i}) \cdot p(y_{i}|\lambda_{i}) d\lambda_{i}}
$$
 (3)

Assuming constant a priori probabilities for the λ_i , the integral in **the denominator of equation (3) is**

$$
\int_{0}^{\infty} p(y_i | \lambda_i) d \lambda_i = \frac{1}{y_i!} \int_{0}^{\infty} \lambda_i^{y_i} \cdot e^{-\lambda_i} d \lambda_i = \frac{1}{y_i!} \cdot \Gamma (y_i + 1) = 1.
$$

Therefore

$$
p(\lambda_{i}|y_{i}) = \frac{\lambda_{i}^{y_{i}}}{y_{i}!} \cdot e^{-\lambda_{i}} = p(y_{i} | \lambda_{i})
$$
 (4)

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The density (4) has a maximum at X. = y. . The measured y. is i $\frac{1}{2}$ **i** $\frac{1}{2}$ *l* $\frac{1}{2}$ **i** $\frac{1}{2}$ *l* $\frac{1}{2}$ *l* therefore the most probable value of λ and for each y_i there belongs **a** distribution in λ_i of the form of equation (4).

If we now attempt to fit a family of curves to n points, each having y. characterized by equation (4), the best fit, by the principle of maximum likelihood, will be obtained when the product P of the probabilities for each point X. is a maximum, i.e . when

$$
P = \frac{n}{\prod_{i=1}^{n} p(\lambda_i | y_i)} = \alpha \text{ maximum.}
$$

This gives

$$
P = \frac{n}{\pi} \frac{\lambda_i^{\gamma_i}}{\gamma_i} \cdot e^{-\lambda_i} = e^{-\sum_{i=1}^{n} \lambda_i} \frac{n}{\pi} \frac{\lambda_i^{\gamma_i}}{\gamma_i!} = a \text{ maximum.}
$$
 (5)

Taking the logarithm this leads to

$$
\ell_{n} p = -\sum_{i=1}^{n} \lambda_{i} + \sum_{i=1}^{n} y_{i} \cdot \ell_{n} \lambda_{i} - \sum_{i=1}^{n} \ell_{n}(y_{i}!) = \alpha \text{ maximum. (6)}
$$

To proceed further we must specify the family of curves which we wish to fit to the data.

4. The form of the alpha line

We write the theoretical line shape (near the high energy edge) for the alpha line in our magnetic spectrograph in the following form (see [2]) .

$$
\lambda_{i} = a(b - x_{i})^{3/2} + y_{o} \t\t(7)
$$

In this equation y_o is the average background and a and b are constants **to be determined. The constant b corresponds to the point on the plate at which the theoretical value X, becomes equal to background, and** it is this quantity which is used to calculate the alpha energy.

The form of equation (7) is chosen in order to avoid subtracting the background from the measured points and taking the 2/3 power of the resulting counts, both of which would distort the Poissonian distribution of the y.. The fit is performed directly to the measured data.

Substituting equation (7) into equation (6) gives

$$
-a \sum_{i=1}^{n} (b-x_i)^{3/2} - n y_0 + \sum_{i=1}^{n} y_i \cdot \ln[a(b-x_i)^{3/2} + y_0] - \sum_{i=1}^{n} \ln(y_i!)
$$

= **a maximum . (8)**

Differentiating (8) with respect to a and b and setting the derivatives equal to zero, we obtain the two simultaneous equations

$$
\sum_{i=1}^{n} (b-x_i)^{3/2} = \sum_{i=1}^{n} \frac{y_i (b-x_i)^{3/2}}{a(b-x_i)^{3/2} + y_o}
$$
 (9a)

and

$$
\sum_{i=1}^{n} (b-x_i)^{1/2} = \sum_{i=1}^{n} \frac{y_i (b-x_i)^{1/2}}{a(b-x_i)^{3/2} + y_0}.
$$
 (9b)

Equatio-, (9a) is obtained by differentiating (8) with respect to a and it is therefore the condition which, when fulfilled, gives the best value of a for b constant. Similarly, equation (9b) is obtained by differentiating (8) with respect to b, and its solution provides the best value of b for a constant. Unfortunately, there is no exact analytic solution, but equations (9) are readily solved numerically by means of a computer. One starts with trial values of a and b and by successive iterations a solution can be obtained to any desi'ed degree of accuracy.

5. Error estimation

Having obtained the best value of the intercept $b = b_0$ from (9), **we would like to estimate its error due to statistical fluctuations in the** data. Unfortunately, equations (9) prove just as resistant to solution for σ_h^2 as they do for **b** itself. However, the error in **b** may be estimated graphically by plotting distribution (5) as a function of b. The procedure is to fix a value of b, find the best value of a using equation (9a), calculate the set of λ_i using (7) and then calculate P from (5). Again, this is easily done with a computer. The resultant probability density will **this is easily done with a computer. The resultant probability density will h** \mathbf{b} \mathbf{c} \mathbf{b} \mathbf{c} \mathbf{d} \mathbf{c} \mathbf{d} \mathbf{c} \mathbf{d} \mathbf{d} probability at the maximum. For a normal distribution this would corr**e**spond to one standard deviation. These upper and lower error estimates will normally be different, reflecting the asymmetry of the initial Poisson **normally be different, reflecting the asymmetry of the initial Poisson**

6. Example

This method has been applied to the analysis of a number of plates from the alpha-particle spectrograph. As an example Fig. 1 shows the high energy edge of the alpha line obtained from a source of ²⁴⁰Pu. The **solid line is the calculated fit to the data. Fig. 2 shows the calculated asymmetric distribution in the intercept with most probable value** $(89.918 + \frac{0.024}{0.014})$ mm.

240 The corresponding energies from an analysis of four plates of Pu are summarized in Table 2 .

Although on theoretical grounds, as explained above, one should expect this method to give more reliable results, extensive numerical calculations with data from our alpha spectrograph show no significant difference between the results for the energies based on this and the usual least-squares method.

 \mathcal{A}^{max}

Fig. 1. High energy edge of an alpha line with an adjusted fit

Fig. 2. Calculated probability distribution for the extrapolated end point b_0 of an alpha line

240 Table 2 . Experimental values for the alpha energy of Pu

For the least-squares calculation the mean and standard deviations were calculated applying the usual formulae

$$
\bar{x} = \frac{\sum x_i/\sigma_i^2}{\sum 1/\sigma_i^2}, \qquad \frac{1}{\sigma_{\bar{x}}^2} = \sum \frac{1}{\sigma_i^2}
$$

For the maximum-likelihood calculation, an approximation was used as explained in the Appendix. The asymmetry, clecrly visible in the errors of the individual measurements, tends to disappear in the mean, as might be expected on the basis of the central limit theorem.

It is certainly reassuring that the two methods of calculation do not give, widely disparate results and one can thei ;fore, for most applications at least, apply the simpler least-squares method with confidence.

 $\mathcal{L}^{\text{max}}_{\text{max}}$

 $\label{eq:2} \frac{1}{2}\sum_{i=1}^n\frac{1}{2}\sum_{j=1}^n\frac{1}{2}\sum_{j=1}^n\frac{1}{2}\sum_{j=1}^n\frac{1}{2}\sum_{j=1}^n\frac{1}{2}\sum_{j=1}^n\frac{1}{2}\sum_{j=1}^n\frac{1}{2}\sum_{j=1}^n\frac{1}{2}\sum_{j=1}^n\frac{1}{2}\sum_{j=1}^n\frac{1}{2}\sum_{j=1}^n\frac{1}{2}\sum_{j=1}^n\frac{1}{2}\sum_{j=1}^n\frac{1}{2}\sum_{j=1}^n\frac{$

Appendix

An approximate method for determining the "best" value from a set of measurements with asymmetric errors

Let us suppose that as a result of a series of n independent measurements on a quantity x we obtain the values x₁, x₂, ..., x_n with **I & n probability distributions f,(x), f»(x), ... , f (x). The point x. is that value of x for which f.(x) is a maximum. We wish to use the measured x. i i to obtain a "best" value x. This "best" value, in the sense of mcximum likelihood, will be the value of x for which the function**

$$
F(x) = f_1(x) \cdot f_2(x) \cdot \ldots \cdot f_n(x) = \frac{n}{i-1} f_i(x)
$$
 (A1)

is a maximum.

In the case we are considering the distributions $f(x)$ have the form **shown in Fig. 2 and must be evaluated numerically. In order to calculate F(x) exactly we would have to determine numerically all the densities f.(x) at a large number of points and then form their product. Although in principle this could be done, we have preferred to simplify the calculation by replacing the f.(x) by a "double" or "asymmetric Gaussian" of the following normalized form.**

$$
g_{i}(x) = \begin{cases} C_{i} \cdot \exp\left[-\frac{(x - x_{i})^{2}}{2 \sigma_{i2}^{2}}\right], & \text{for } -\infty < x \leq x_{i} \\ C_{i} \cdot \exp\left[-\frac{(x - x_{i})^{2}}{2 \sigma_{i1}^{2}}\right], & \text{for } x_{i} \leq x < \infty, \end{cases}
$$
\nwith C_i = $\sqrt{\frac{2}{\pi}} \cdot (\sigma_{i1} + \sigma_{i2})^{-1}$.

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Here x. is the most probable value (b in Fig. 2) and <j-.. and i o il σ_{12}^- are the upper and lower confidence limits, respectively, as **described in section 5 and shown in Fig. 2 .**

It should be noted that as a consequence of the asymmetry of equation (A2), the mean value of x

$$
E_{i}(x) = \int_{-\infty}^{\infty} x \cdot g_{i}(x) dx = x_{i} + \sqrt{\frac{2}{\pi}} \cdot (\sigma_{i1} - \sigma_{i2})
$$
 (A3)

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cannot be identical with x_1 , unless σ_{11} = σ_{12} .

The "best" value \bar{x} of x is therefore that for which

$$
F(x) = \frac{n}{\pi} g_i(x) \text{ is a maximum.}
$$

As before, the upper and lower confidence limits are the points x at which

 $F(x) \approx 0.607 F(\bar{x})$.

We may add that the variance of the "asymmetric Gaussian" (A2) is given **by**

$$
E_i(x - x_i)^2 = \sigma_{i1} \cdot \sigma_{i2} + (1 - \frac{2}{\pi})(\sigma_{i1} - \sigma_{i2})^2. \tag{A4}
$$

This value can be shown to lie between σ_{i1}^2 and σ_{i2}^2 .

References

[1] A. Rytz: Helv. Phys. Acta 34, 240 (1961)

[2] B. Grennberg, A. Rytz: Métrologie 7, 65 (1971)

f_3J B.L. van der Waerden: Mathematical Statistics (Springer, Berlin, 1969)

£4J J.W. MUller: Traitement statistique des résultats de mesure, Rapport BIPM-108 (1969 ff).

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