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OSCILLATION AND MIXING OF MASSLESS ELECTRON AND MUON NEUTRINOS

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Abstract

It is not necessary that the electron neutrino and muon neutrino themselves have mass for v_{e} - v_{μ} oscillations to occur. For N neutrinos, if one is massive, oscillations occur with unique oscillation length determined by the mass of the heavy neutrino, mimicing two neutrino mixing. Any observation of v_{e} - v_{μ} oscillations cannot therefore be regarded as evidence for a finite mass v_e or v_u .

The conventional discussion of $v_a - v_a$ oscillations assumes that at **e u least one of these two neutrinos is massive. The discovery of a third** generation neutrino, v_{τ} , and the possibility that there may be additional **2 3 neutrinos makes more complex oscillation phenomena possible. ' It does not** appear to have been noticed that in these circumstances $v_{\alpha} - v_{\alpha}$ oscillations ?. t! can occur even if both v_e and v_u are massless. All that is required is that **at least one of the neutrinos has a finite mass, and that it is kinematically possible for this heavy neutrino to be produced in the reaction producing the neutrino beam. An immediate consequence of this observation is that** observation of $v_a - v_a$ oscillations does not imply limits on the masses of v_{e} v_{u} , contrary to popular belief, and that a reanalysis of the existing **experimental limit is necessary.**

2 We consider first the general neutrino system. The eigenstates of the mass matrix will be denoted by v_1 (i = 1, ..., n) and the neutrino states **which diagonalise the weak interaction currents by N^. The corresponding charged leptons will be labelled L..**

$$
N_{i} = \sum_{j} U_{i,j} \vee_{j} ; \vee_{j} = \sum_{j} N_{i} U_{i,j}^{\star}
$$
 (1)

is the unitary transformation connecting the weak interaction and mass eigenstates.

Suppose we start with the weak interaction production of N^2 with **momentum P at time t = 0. Note that it must be kinematically possible to** produce all of the v_1 in the reaction which provides the beam for N^2 to be **the initial state. The neutrino state at a later time t is**

$$
|2; t \rangle = e^{-iHt} |N_2 \rangle
$$

= $\sum_{j} U_{2j} e^{-iE_j t} |V_j \rangle$ (2)

- 2 - *?*

 E_i = m_i ² + P² where m_i is the mass of the jth neutrino. The probability of observing the weak production of the charged lepton L_1 at the time t is proportional to the probability ${\tt P}_{2 \, {\bf l}} \, ({\tt t})$ that the neutrino is in the state N_1 at this time

$$
P_{21}(t) = |\langle N_1 | 2, t \rangle|^2
$$

= $\|\sum_{j} U_{1j}^* U_{2j} e^{-1E_j t}\|^2$ (3)

Notice that if all the neutrinos are degenerate in mass $P_{21}(t) = 0$. This confirms the fact that in these circumstances the neutrino labeling can be defined by the N_1 , as these are also mass eigenstates.

In general if at least one neutrino has a different mass to the other equation (3) will contain oscillatory terms. This special neutrino need not be either v_1 or v_2 . Since a detailed analysis of (3) would serve to confuse rather than illustrate the main point of this letter, I specialise to the case where $n - 1$ of the neutrinos have zero mass and one, v_n , has a non zero n' mass m. Then $E_j = P$ for $j = 1, ..., n - 1$ and

$$
P_{21}(t) = \left| \sum_{j=1}^{n-1} U_{2j} U_{1j}^* + e^{i\Delta t} U_{2n} U_{1n}^* \right|^2
$$

where $\Lambda = P - E_n$. The unitarity of U shows that

n

$$
\sum_{j=1}^{n-1} U_{2j} U_{1j}^* = -U_{2n} U_{1n}^*
$$

so in this case

$$
P_{21}(t) = |2 U_{2n} U_{1n}^*|^2 \sin^2 \frac{\Delta t}{2}
$$
 (4)

4 This reduces to the usual formula for two neutrino mixing on setting $n = 2$, $U_{22} = \cos \theta$, $U_{12} = \sin \theta$. This suggests we define θ_{12}^* , an effective $mixing angle, by sin 2\theta_{12}^* = |2 U_{2n}U_{1n}^*|$

Then
$$
P_{21}(t) = \sin^2 2\theta_{12}^* \sin^2 \frac{\Delta t}{2}
$$
, (5)

which has the same form as the usual formula.

A similar calculation shows that P_{22} (t), the probability of detecting **type 2 neutrinos in the beam at time t may also be cast in the usual form**

$$
P_{22}(t) = 1 - \sin^2 2\theta_2 \frac{\pi}{2} \sin^2 \frac{\Delta t}{2}
$$
 (6)

with a new effective mixing angle defined by

$$
\sin^2 2\theta_{22}^{\star} = 4 U_{2n} U_{2n}^{\star} (1 - U_{2n} U_{2n}^{\star})
$$

$$
(1.e. either cos2 \theta22 = U2n U2n* or sin2 \theta22 = U2n U2n*).
$$

As long as oscillations are detected by observing one type of neutrino only, the present model mimics the standard massive $v_e - v_\mu$ mixing **model. It can be distinguished if more than one type of neutrino is detected, since the effective mixing angles are not the same.**

Even if a sinusoidal $v_{e} - v_{\mu}$ oscillation with a single period is **detected, this phenomenon is not evidence for a v - v mass difference. The e p** finite mass. In the general case the oscillation will be a complicated mixture of a number of different oscillation periods.

In the case $m \ll P$, the oscillation time Δ^{-1} becomes $2P/m^2$ and (5)

takes the form

$$
P_{21}(t) = \sin^2 2\theta_1 \frac{\pi}{2} \sin^2 \frac{\pi^2 t}{4P}
$$
 (7)

This formula, with the interpretations $\theta_{12}^* = \theta$, the v_{e^-} . mixing angle, and $m = \nu m_1^2 - m_2^2 = M_{1/2}$, was used to analyse the attempt to observe $v_1 - v_2$ **u e ue'** *J r \i e* oscillations in the CERN neutrino beam. The results of that experiment were μ expressed as allowed regions in the μ is well represented by the analytic expression

$$
\sin 2\theta M_{\mu e}^2 < 1 eV^2.
$$

Our analysis shows that the data can equally well be regarded as placing restrictions on the mass and mixing angles of a single heavy neutrino,

$$
2 U_{2n} U_{1n}^* m^2 < 1 eV^2
$$

and that more complex interpretations are possible.

Clearly observation of neutrino oscillations, while it will demonstrate the existence of at least one massive neutrino will not provide us with unambiguous information on the neutrino masses. We will not even know which neutrino has the mass.

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References

- **1. B. Pontecorvo, Soviet Phys. JETP 5_3 1717 (1967)**
- **2. N. Cabibbo, Phys. Lett.** *J2J&* **333 (1979)**
- **3. L. Wolfensten, Carnegie Mellon University Preprint (1979)**
- **4. J. Blietschen et al, Nucl. Phys. B133 205 (1978)**

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