# Appendix to "Periodograms for movement ecologists: a non-parametric method to uncover periodic patterns of space use in animal tracking data"

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# <sup>1</sup> Appendix A: FFT Lomb-Scargle periodogram

2 A periodogram estimates the spectral-density function  $\tilde{\sigma}(f)$  at frequency f, which is itself the

3 Fourier transform of the stationary (or time-averaged) autocorrelation function 
$$
\sigma(\tau)
$$
.

$$
\tilde{\sigma}(f) \equiv \int_{-\infty}^{+\infty} dt \, e^{-2\pi i f \tau} \, \sigma(\tau) \propto \text{VAR} \left[ \int_{-\infty}^{+\infty} dt \, e^{-2\pi i f t} \, x(t) \right] = \text{VAR}[\tilde{x}(f)] \;, \tag{A.1}
$$

$$
\overline{5}
$$

 $\sigma(\tau) = \lim_{T \to \infty}$ 1 2T  $\int_{0}^{+T}$  $-T$ dt  $\overline{ \text{time average}}$ time average  $\sigma(\tau) = \lim_{\epsilon \to 0} \frac{1}{2\pi} \int dt \text{COV}[x(t + \tau), x(t)],$ (A.2)

6

<sup>7</sup> where the noted time average is unnecessary for process with stationary autocorrelation func- $\epsilon$  tions that have no dependence upon absolute time t, and where we define the variance of any of complex-valued process  $VAR[z] \equiv \langle z z^* \rangle$ , where  $z^*$  denotes the complex conjugate of z. The DFT

<sup>10</sup> periodogram is a straightforward approximation to the variance of the (continuous) Fourier trans-11 form of  $x(t)$  via the discrete Fourier transform (DFT). Another way of motivating the periodogram, 12 which more readily generalizes in the case of missing data, is that the periodogram at frequency  $f$ 13 is equivalent to the the 'power' derived from a least-squares fit of all  $n$  data points to a sinusoid  $14$  of frequency f. "Power" in this context is bit of a mathematical abstraction that we will later 15 define, but note that if the signal  $x(t)$  represents oscillator position, velocity, electrical current, or 16 electrical voltage, then the signal's variance  $\sigma(0)$  is proportional to its average oscillator potential, <sup>17</sup> kinetic, inductor, or resistor energy, respectively. Further note that with Parseval's theorem we can <sup>18</sup> distribute this "energy" among times or frequencies

$$
\sigma(0) \propto \int_{-\infty}^{+\infty} dt \, \text{VAR}[x(t)] = \int_{-\infty}^{+\infty} df \, \text{VAR}[\tilde{x}(f)]. \tag{A.3}
$$

<sup>21</sup> Therefore the periodogram estimates autocorrelation structure via the spectral-density function and <sup>22</sup> distributes the signal's variance among frequencies in a way that is conjugate to how the variance <sup>23</sup> is distributed among times.

24 The naive implementation of the DFT or LS periodogram has a computational cost of  $\mathcal{O}(n^2)$ 25 to estimate the n most relevant frequencies within the range  $df \leq f \leq F$  (Table [A.1\)](#page-2-0), where df is the natural frequency resolution of the data and F is the Nyquist frequency or natural frequency 27 cutoff of the data. The fast Fourier transform (FFT) reduces this computational cost to  $\mathcal{O}(n \log n)$  for the DFT periodogram. Furthermore, [Press and Rybicki](#page-11-0) [\(1989\)](#page-11-0) discovered that, after expanding the sinusoids in an evenly spaced grid via Lagrange interpolating polynomials, the LS periodogram can also be calculated using FFT techniques. As the Lagrange interpolants are approximations of the true functions, the evenly-spaced grid needs to be fine enough to accurately capture the high-frequency behavior. Our implementation of the LS periodogram differs in that it is given by a simple expression without the necessity of Lagrange interpolation. In the case of evenly scheduled data, where our implementation is exact, the cost savings of avoiding Lagrange interpolation are <sup>35</sup> typically on the order of a factor of 10-40 [\(Press and Rybicki,](#page-11-0) [1989\)](#page-11-0).



<span id="page-2-0"></span>Table A.1: Conjugate relationship between the temporal quality of data and the frequential quality of data under the DFT. These relations are still approximately true for the Lomb–Scargle periodogram (LSP).

36 For our derivation, we exploit the fact that for the realization of the process of interest  $x(t)$ , we <sup>37</sup> know the indicator function

$$
w(t) \equiv \begin{cases} 1, & x(t) \text{ observed} \\ 0, & x(t) \text{ missed} \end{cases}, \tag{A.4}
$$

40 and, though we have not always measured  $x(t)$  on a uniform time grid, we have measured  $w(t)$ 41 and (effectively)  $w(t)x(t)$  on a uniform time grid. This scheme is natural for data that are evenly <sup>42</sup> scheduled but feature missing values, as the observed times neatly reside in a uniformly spaced grid <sup>43</sup> of scheduled observation times that is only slightly larger than the number of observations.

<sup>44</sup> Our fast implementation of the LS periodogram then follows the derivation of [Scargle](#page-11-1) [\(1982\)](#page-11-1). <sup>45</sup> We exploit the equivalence of the LS periodogram to a least-squares fit of the data to independent <sup>46</sup> sinusoids

<span id="page-2-2"></span>
$$
x(t) \approx x_f(t) = A(f) e^{+2\pi i f t} + A(f)^* e^{-2\pi i f t}, \tag{A.5}
$$

49 to obtain the amplitudes  $A(f)$ . This fitting has the associated least-squares cost function

<span id="page-2-1"></span>
$$
L(f) = \sum_{t} w(t) |x(t) - x_f(t)|^2, \tag{A.6}
$$

 $52$  given weights  $w(t)$ . Solutions for the amplitudes must therefore satisfy the system of equations

$$
\sum_{t=54}^{53} \sum_{t=54} w(t) \sum_{t=54} e^{-4\pi i f t} w(t) \begin{bmatrix} \hat{A}(f) \\ \hat{A}(f)^* \end{bmatrix} = \begin{bmatrix} \sum_{t=54}^{53} e^{-2\pi i f t} w(t) x(t) \\ \sum_{t=54}^{54} e^{-2\pi i f t} w(t) x(t) \end{bmatrix} . \tag{A.7}
$$

55 If we represent the data  $x(t)$  and weights  $w(t)$  on a uniform time grid, with missing data naturally <sup>56</sup> weighted by zero, then all of these sums can be calculated with the FFT implementation of the 57 DFT of the weights  $w(t)$ 

$$
\text{DFT}\{w\}(f) = W(f) = \sum_{t} e^{-2\pi i f t} w(t), \tag{A.8}
$$

<sup>60</sup> so that we have

<span id="page-3-0"></span>
$$
\begin{bmatrix} W(0) & W(2f) \\ W(2f)^* & W(0) \end{bmatrix} \begin{bmatrix} \hat{A}(f) \\ \hat{A}(f)^* \end{bmatrix} = \begin{bmatrix} \text{DFT}\{wx\}(f) \\ \text{DFT}\{wx\}(f)^* \end{bmatrix} . \tag{A.9}
$$

63 For frequencies f larger than half the Nyquist frequency  $F$ , it is convenient to exploit the periodicity <sup>64</sup> of the transform

$$
W(2f) = W(2f - 2F). \tag{A.10}
$$

 $67$  Solving Eq.  $(A.9)$  for the amplitudes, we then have

<span id="page-3-2"></span>
$$
\hat{A}(f) = \frac{W(0)\,\text{DFT}\{wx\}(f) - W(2f)\,\text{DFT}\{wx\}(f)^*}{W(0)^2 - |W(2f)|^2},\tag{A.11}
$$

 $\sigma$  while the power estimate in [Scargle](#page-11-1) [\(1982\)](#page-11-1) Eq. (C3) equates to

<span id="page-3-1"></span>
$$
\hat{P}(f) = \frac{1}{2} \left( \sum_{t} w(t) |x(t)|^2 - \min_{A(f)} L(f) \right), \tag{A.12}
$$

<sup>73</sup> by expressing his power estimate in terms of both  $x(t)$  and  $w(t)$ . Refer to [Scargle](#page-11-1) [\(1982\)](#page-11-1) for more <sup>74</sup> details on the "power", but note simply that it measures the amount of variability in the data <sup>75</sup> attributable to frequency f. Combining equations  $(A.12)$ ,  $(A.6)$ , and  $(A.5)$ , we now have

$$
\hat{P}_{\text{LSP}}(f) = -\sum_{t} w(t) x(t) \hat{x}_f(t) + \frac{1}{2} \sum_{t} w(t) \hat{x}_f(t)^2, \tag{A.13}
$$

$$
= -\hat{A}(f) \sum_{t} e^{+2\pi i f t} w(t) x(t) - \hat{A}(f)^{*} \sum_{t} e^{-2\pi i f t} w(t) x(t)
$$
 (A.14)

$$
+\frac{1}{2}\hat{A}(f)^2\sum_{t}e^{+4\pi ift}w(t)+\left|\hat{A}(f)\right|^2\sum_{t}w(t)+\frac{1}{2}\hat{A}(f)^{2*}\sum_{t}e^{-4\pi ift}w(t)\,,
$$

$$
= -\hat{A}(f) \, \text{DFT}\{wx\}(f)^* - \hat{A}(f)^* \text{DFT}\{wx\}(f) \tag{A.15}
$$

$$
+\frac{1}{2}\hat{A}(f)^2 W(2f)^* + |\hat{A}(f)|^2 W(0) + \frac{1}{2}\hat{A}(f)^{2*} W(2f) ,
$$

 $\mathfrak{g}_2$  and then combining this expression with Eq.  $(A.11)$  we have after some simplification

<span id="page-4-0"></span>
$$
\hat{P}_{\text{LSP}}(f) = \frac{W(0) \left| \text{DFT}\{wx\}(f) \right|^2 - \text{Re}\left[W(2f)^* \text{DFT}\{wx\}(f)^2\right]}{W(0)^2 - |W(2f)|^2},\tag{A.16}
$$

85 which can be constructed entirely from the FFT of  $w(t)$  and  $w(t) x(t)$ .

#### <sup>86</sup> A.1 Comparison to the DFT periodogram

 Strictly speaking the DFT periodogram is not defined for missing data, but researchers not familiar with the Lomb-Scargle periodogram will typically replace a few missing values of the data with the mean or an interpolated value. In the case of interpolating the missing value, the effect is to bias the autocorrelation estimate with the properties of the interpolating function at the scale of the 91 gaps. I.e., straight-line interpolation over gaps of width  $\Delta t$  gives the appearance of a more ballistic 92 process over timescales  $\leq \Delta t$  and frequency scales of  $\geq 1/\Delta t$ . On the other hand, the case of replacing missing data with the mean value can be derived from a simpler cost function than that

 of the Lomb-Scargle periodogram, where relations between different phase sinusoids of the same frequency are ignored. With the mean first detrended from the data, then the single-frequency fitting and cost functions are given by

97 
$$
x(t) \approx x_f(t) = A(f) e^{+2\pi i f t}
$$
,  $L(f) = \sum_t w(t) |x(t) - x_f(t)|^2$ , (A.17)

<span id="page-5-0"></span>with the amplitude solutions and power estimates found to be

$$
\hat{A}(f) = \frac{\text{DFT}\{wx\}(f)}{W(0)}, \qquad \hat{P}_{\text{DFT}}(f) = \frac{|\text{DFT}\{wx\}(f)|^2}{W(0)}, \qquad (A.18)
$$

102 where again  $W(f) = DFT\{w\}(f)$  and  $W(0) = n$ . Comparing this result to the Lomb-Scargle 103 periodogram relation that we have derived  $(A.16)$ , we can see that the mean-imputed DFT peri- odogram shows up in the LSP, but it is shifted and rescaled by terms that vanish when  $W(2f) = 0$ . 105 If  $w(t) = 1$  for all evenly-sampled t, then  $W(f) = 0$  for all canonical  $f \neq 0$  and so the LSP reduces to the DFT periodogram. In other words, if there are no missing data, the LSP and DFT are strictly equivalent. Note, however, that the ordinary method of inflating the frequency resolution of the DFT periodogram by padding the data with mean values is also not consistent with the LSP value, as padding the data produces a periodogram equivalent to Eq. [\(A.18\)](#page-5-0). In short, the LSP gives an improved result both when there are missing data and when the frequency resolution is inflated beyond the natural resolution of the data.

## A.2 Expectation value of the periodogram

 As noted by [Scargle](#page-11-1) [\(1982\)](#page-11-1), the DFT and LS periodograms tend to be fairly similar in practice, and 114 so one can view the LSP in Eq. [\(A.16\)](#page-4-0) as being the DFT periodogram  $|DFT{wx}(f)|^2/W(0)$  with small corrections. Here we investigate this dominant term of the Lomb-Scargle periodogram under different sampling regimes. For simplification we will assume that the sampling and movement

117 processes are independent,  $\langle w(t) x(t') \rangle = \langle w(t) \rangle \langle x(t') \rangle$ , and that the movement process is stationary 118 with detrended mean and autocorrelation function  $\text{COV}[x(t), x(t')] = \sigma(t-t')$ . In this case we can <sup>119</sup> express the expectation value

$$
\left\langle \hat{P}_{\text{LSP}}(f) \right\rangle \approx \frac{1}{W(0)} \left\langle |\text{DFT}\{wx\}(f)|^2 \right\rangle, \tag{A.19}
$$

<span id="page-6-1"></span>
$$
\approx \frac{1}{W(0)} \sum_{tt'} e^{-2\pi i f(t-t')} \left\langle w(t) w(t') \right\rangle \sigma(t-t'). \tag{A.20}
$$

123 For the next step we perform an inverse Fourier transform from the frequency  $f$  domain back to 124 the time-lag  $\tau$  domain to obtain the equivalent relation

$$
\text{DFT}^{-1}\Big\{\Big\langle\hat{P}_{\text{LSP}}\Big\rangle\Big\}(\tau) \approx \frac{1}{W(0)} \sum_{ftt'} e^{+2\pi\imath f(\tau - [t-t'])} \langle w(t) w(t')\rangle \sigma(t-t'). \tag{A.21}
$$

127 If the data vectors  $x(t)$ ,  $w(t)$  are padded with zeros to twice their recorded length 2N (on a uniform <sup>128</sup> grid), then the frequency sums evaluate to Kronecker delta functions with the identity

$$
\delta_K(t) = \frac{\sum_f e^{+2\pi i f t}}{\sum_f} \tag{A.22}
$$

<sup>131</sup> and so we have

$$
\text{DFT}^{-1}\left\{ \left\langle \hat{P}_{\text{LSP}} \right\rangle \right\} (\tau) \approx \frac{2N}{W(0)} \sum_{tt'} \delta_{\text{K}} (\tau - [t - t']) \left\langle w(t) \, w(t') \right\rangle \sigma(t - t'), \tag{A.23}
$$

<span id="page-6-0"></span>
$$
\approx \frac{2N}{n(0)} n(\tau) \sigma(\tau), \qquad (A.24)
$$

 using the property that the Kronecker delta function is equal to one when its argument is zero and 136 zero otherwise, and where  $n(\tau)$  is the number of data pairs recorded with time-lag  $\tau$  between them. If the data vectors were not padded with zeros to at least twice their recorded length  $2N$ , and were left with length N, then the second half of the estimate would be a repeat of the first half due to the periodicity of this representation of the Kronecker delta function.

140 The pair number  $n(\tau)$  is also proportional to the inverse transform of the periodogram of the sampling process  $w(t)$ . Therefore, back in the frequency domain, the periodogram of the movement process actually represents a convolution of the autocorrelation functions of the sampling process and the movement process. One might imagine de-biasing this estimator by dividing the result 144 by a factor of  $n(\tau)$  in the time domain. However, the resulting correlogram would no longer be positive definite, which renders the corresponding periodogram illegitimate in approximating the strictly positive eigenvalue spectrum of a positive-definite covariance matrix.

#### 147 A.3 Autocorrelation in the sampling schedule

148 Equation [\(A.24\)](#page-6-0) shows the direct link between the autocorrelation function  $\sigma(\tau)$  and the estimated periodogram  $\tilde{P}(f)$ , as given by the number of observation pairs  $n(\tau)$  with time lag  $\tau$  between 150 them. Importantly, the periodogram is always biased by the sampling schedule through  $n(\tau)$ . 151 When there are no missing data,  $n(\tau)$  is a simple function that decreases linearly with the time 152 lag  $\tau$ , and its influence on the periodogram is both predictable and mild. If there are missing data with no particular autocorrelation structure in timing of their gaps (uncorrelated and uniformly 154 distributed), then again  $n(\tau)$  has a simple structure that does not contaminate the periodogram in a non-trivial way. However, if the sampling schedule is itself periodic, or otherwise temporally autocorrelated, then both the DFT and LS periodograms will exhibit strong biases that say more about the autocorrelation structure of the sampling schedule than that of the signal. This is why the diagnostic argument of ctmm's plot.periodogram method will plot both the periodogram of the data and the periodogram of the sampling schedule. Viewing both periodograms side-by-side allows users to check for any possible autocorrelation structure in the sampling schedule that might have propagated into the periodogram of the data.

#### A.4 Sampling properties and periodogram averaging

 The sampling regime intuitively constrains the resolution and bandwidth of the periodogram—i.e., the precision and range of frequencies over which the period can be estimated. Although, the Lomb-Scargle periodogram can be calculated for any frequency, it only contains novel information over some limited set of frequencies. In the case of evenly sampled data where the LSP reduces to the DFT periodogram, this relationship is exact and is summarized by table [A.1.](#page-2-0) In short, temporal 168 resolution translates into frequency range, via the Nyquist frequency  $F = 1/(2dt)$ , while temporal range translates into frequency resolution. With the canonical set of frequencies described in table [A.1,](#page-2-0) the periodogram contains approximately two locations worth of information per positive frequency. By default, the ctmm periodogram function inflates this resolution by a factor of two with the res=1 argument, to make this relationship approximately one-to-one. Inflating the frequency rrs resolution further will cause the estimates  $\hat{P}(f)$  to be increasingly correlated between frequencies and lead to a locally smooth periodogram. On the other hand, if the periodogram is evaluated at frequencies beyond the Nyquist frequency, then the periodogram simply repeats itself, as there is no information in the data beyond this cutoff. The Lomb-Scargle periodogram approximately follows the same general relations as the DFT periodogram, though some frequencies can be better sampled than others.

 Autocorrelation estimates of any kind can be very noisy for an individual and pooling the esti- mates of multiple individuals is a good way of reducing this variability. An assumption required to average periodograms is that the sampling frequencies are roughly the same, as differently struc- tured sampling schedules lead to different natural biases of the periodogram. To average multiple 183 individual periodograms, we choose the best frequency resolution (largest  $T$ ), worst Nyquist fre- quency (largest dt), and then weight the estimates by their corresponding amount of data. More 185 specifically, we distribute the n degrees of freedom in a dataset over what would be the natural set of 186 frequencies  $0 < f \leq F$  for that dataset, regardless of how we fix those parameters for the population  estimate. The worst Nyquist frequency is chosen because evaluating any one periodogram beyond its Nyquist frequency yields a nonsense estimate contaminated by Nyquist frequency periodicities, while the best frequency resolution is chosen because inflating a low-resolution periodogram will only induce correlated errors and we account for this in the average via weighting. This selection criteria is automated by the ctmm periodogram function when feeding it a list of telemetry objects.

## 192 A.5 How periodicity combines with random motion

 In App. D we explore periodograms for all of the basic continuous-time movement models, including Brownian motion and Ornstein–Uhlenbeck motion, without any periodicity. These basic movement models provide what a signal analyst might refer to as nuisance "background noise" in which the periodic "signal" exists, in that, for a Gaussian stochastic process the expectation value of the periodogram decomposes into

$$
\langle \hat{P}(f) \rangle = \langle \hat{P}_{\text{stochastic}}(f) \rangle + \langle \hat{P}_{\text{deterministic}}(f) \rangle, \tag{A.25}
$$

where the stochasic component describes an (on average) smooth curve (e.g.,  $1/f<sup>2</sup>$  for Brownian motion) and the deterministic mean component gives us the "peak" or "spike" atop this curve. Therefore the importance of the height of any periodicity is relative to the background curve of a model such as Brownian motion.

#### A.6 Effect of telemetry error on the periodogram

 For additive telemetry errors that are uncorrelated with the movement process, the expectation value of the periodogram of the noisy, observed process  $\hat{P}_{data}(f)$  is given by the sum

$$
\langle \hat{P}_{\text{data}}(f) \rangle = \langle \hat{P}_{\text{move}}(f) \rangle + \langle \hat{P}_{\text{error}}(f) \rangle \,, \tag{A.26}
$$

<sup>209</sup> of the animal's movement process and the error process. The autocorrelation function of an un-210 correlated error process  $\sigma_{\rm error}(\tau)$  is furthermore proportional to a Dirac delta function  $\delta(\tau)$ , and so 211 from Eq. [\(A.20\)](#page-6-1) the quantity  $\langle \hat{P}_{error}(f) \rangle$  is approximately constant for all frequencies f. Therefore, <sup>212</sup> the effect of telemetry error is to shift the resulting periodogram vertically along the y-axis and <sup>213</sup> induce further variability in the estimate.

## <sup>214</sup> A.7 A gridding algorithm

 To apply the periodogram, we want to define a regular temporal grid that is evenly spaced and can accommodate all data points. As GPS fixes can be delayed, there can be variability in the realized sampling intervals of tracking data. To construct a well behaved grid that avoids an unnecessary temporal resolution, we optimize the alignment of our temporal grid relative to the data. Given the data, a regular temporal grid (sampling schedule) can be defined by two parameters: an initial 220 time  $t_0$  and a grid spacing  $\Delta t$ . For the grid spacing  $\Delta t$ , by default, in ctmm we use the median realized sampling interval, which performs well if there is a single intended sampling rate. For the initial time, we minimize the cost function

$$
COST(t_0) = \sum_{i=1}^{n} \left| \sin\left(\pi \frac{t_i - t_0}{\Delta t}\right) \right|^2,
$$
\n(A.27)

 which is zero if all recorded times are aligned with the grid and greater than zero otherwise. This 226 cost function has the necessary features of being periodic in the parameter  $t_0$  (with period  $\Delta t$ ) and increasing monotonically with increasing misalignment. Moreover, this cost function is both analytic and exactly solvable. Differentiating and expanding our cost function, we then have the optimal grid relation

<span id="page-10-0"></span>
$$
0 = \langle S \rangle \cos \left( 2\pi \frac{\hat{t}_0}{\Delta t} \right) + \langle C \rangle \sin \left( 2\pi \frac{\hat{t}_0}{\Delta t} \right), \tag{A.28}
$$

<sup>232</sup> in terms of the averages

$$
\langle S \rangle = \frac{1}{n} \sum_{i=1}^{n} \sin \left( 2\pi \frac{t_i}{\Delta t} \right), \qquad \qquad \langle C \rangle = \frac{1}{n} \sum_{i=1}^{n} \cos \left( 2\pi \frac{t_i}{\Delta t} \right). \tag{A.29}
$$

235 It then derives from Eq. $(A.28)$  the following formula for the optimal initial time

$$
\hat{t}_0 = -\frac{\Delta t}{2\pi} \tan^{-1} \left( \frac{\langle S \rangle}{\langle C \rangle} \right). \tag{A.30}
$$

<sup>238</sup> This formula is implemented in ctmm to propose a default time grid upon which to compute the <sup>239</sup> LSP.

# <sup>240</sup> References

- <span id="page-11-0"></span><sup>241</sup> Press, W. H., and G. B. Rybicki. 1989. Fast algorithm for spectral analysis of unevenly sampled <sup>242</sup> data. Astrophysical Journal 338:277–280.
- <span id="page-11-1"></span><sup>243</sup> Scargle, J. D. 1982. Studies in astronomical time series analysis. II - Statistical aspects of spectral <sup>244</sup> analysis of unevenly spaced data. The Astrophysical Journal 263:835–853.